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**Aerosol optical properties at
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Mediterranean) – 1.**

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Aerosol optical properties at Lampedusa (Central Mediterranean) – 1. Influence of transport and identification of different aerosol types

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Aerosol optical depth and Ångström exponent were obtained from multi filter rotating shadowband radiometer (MFRSR) observations carried out at the island of Lampedusa, in the Central Mediterranean, in the period July 2001–September 2003. The average aerosol optical depth at 495.7 nm, τ , is 0.24 ± 0.14 ; the average Ångström exponent, α , is 0.86 ± 0.63 . The observed values of τ range from 0.03 to 1.13, and the values of α vary from -0.32 to 2.05, indicating a large variability in aerosol content and size. In cloud-free conditions, 36% of the airmasses come from Africa, 25% from Central-Eastern Europe, and 19% from Western France, Spain and the North Atlantic. In summer, 42% of the airmasses are of African origin. In almost all cases African aerosols display high values of τ and low values of α , typical of Saharan dust (average values of τ and α are 0.36 and 0.42, respectively). Particles originating from Central-Eastern Europe show relatively large average values of τ and α (0.23 and 1.5, respectively), while particles from Western France, Spain and the North Atlantic show the lowest average values of τ (0.15), and relatively small values of α (0.92). Intermediate values of α are often connected with relatively fast changes of the airmass originating sector, suggesting the contemporary presence of different types of particles in the air column. The largest values of α (about 2) were observed in August 2003, when large scale forest fires in Southern Europe produced consistent amounts of fine combustion particles that were transported to the Central Mediterranean by a persistent high pressure system over Central Europe. Smoke particles in some cases mix with desert dust, producing intermediate values of α . The seasonal distribution of the meteorological patterns over the Mediterranean, the efficiency of the aerosol production mechanisms, and the variability of the particles' residence time produce a distinct seasonal cycle of aerosol optical depths and Ångström exponent values. Particles originating from all sectors show a summer maximum in aerosol optical depth. The summer increase in optical depth for European aerosols is linked with an increment in the values of α that indicates an enhancement in the number of fine particles.

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The summer maximum of τ for African particles is associated with a weak reduction in the Ångström exponent, suggesting an increase in the total number of particles and a relatively more intense transport of large particles. The observations were classified according to the aerosol optical properties, and two main classes have been identified: desert dust and biomass burning/urban-industrial aerosols. Values of τ and α averaged over the whole observing period are 0.37 and 0.15 for desert dust, and 0.27 and 1.77 for urban-industrial/biomass burning aerosols. Lampedusa reveals a stronger influence of desert dust compared to other Mediterranean sites (mostly located on the coasts of Europe).

1. Introduction

The Mediterranean sea, the largest enclosed basin of the Earth, is a region characterized by positive average net radiation flux (the amount of incoming solar radiation exceeds the infrared emission) surrounded by areas characterized by negative average net fluxes (Bolle, 2003). The continents surrounding the Mediterranean basin have different surface characteristics, and a complex orography. The radiative equilibrium of the basin is thus controlled by a large number of factors, resulting from the mutual influence of the sea and the continents. Aerosols play an important role: they attenuate the solar radiation (e.g. Markowicz et al., 2002; di Sarra et al., 2003), modify its vertical distribution (e.g. Meloni et al., 2003a, b), affect the infrared balance, and, indirectly, influence the cloud structure and properties (e.g. Levin et al., 1996). Due to the variety of the regions around the basin, different classes of particles can be found in the Mediterranean atmosphere: desert dust, originated from the Sahara desert and from arid regions in the Iberian Peninsula; polluted particles, produced mainly in urban and industrial areas of Continental and Eastern Europe; marine aerosol, continuously formed over the Mediterranean itself or transported from the North Atlantic; and biomass burning particles, often produced in forest fires, mainly during the summer. Moreover, the aerosol properties and the atmospheric chemical composition are influenced by long-

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range transport (see e.g. Lelieveld et al., 2002).

The two main quantities describing how aerosols influence the radiative balance are the aerosol optical depth, which is an extensive property related to the amount of particles, and the single scattering albedo, which is an intensive property related to the capability to absorb radiation. Other factors, such as size distribution, composition, shape, and vertical distribution, also play a role. Various aerosol types display different microphysical and optical properties. Optical depth, composition, size distribution, shape (consequently single scattering albedo, phase function and asymmetry factor), and vertical distribution vary largely in the Mediterranean, thus affecting the radiative balance in different ways. For example, it has been shown that the radiative forcing efficiency (i.e., radiative forcing per unit aerosol optical depth) produced by desert dust in the visible spectral range may differ by a factor of 2 at the surface, and by a factor of 3 at the top of the atmosphere (Meloni et al., 2004), depending on the aerosol source region. Different aerosol types produce negative or positive forcings at the top of the atmosphere over the Mediterranean basin (Horvath et al., 2002; Meloni et al., 2003a). Thus, the determination of the aerosol properties is crucial to understand their influence on radiation and climate.

Measurements of aerosol properties over the Mediterranean have been conducted since the 1980s (see the extensive survey by Smirnov et al., 2002), but most of them were intermittent, and lasted only for short periods. Later measurements covered longer time periods (Cachorro et al., 2000; Formenti et al., 2001; Sabbah et al., 2001; Tanré et al., 2001; Gerasopoulos et al., 2003; Israelevich et al., 2003; Esposito et al., 2004); most of these studies are based on observations carried out in the coastal regions of the Mediterranean. Analyses based on satellite observations were also carried out to identify the transport of Saharan dust (e.g. Dulac et al., 1992; Moulin et al., 1998; Israelevich et al., 2002; Jamet et al., 2004). Systematic measurements of the aerosol distribution and properties are particularly useful in the open Mediterranean sea, because of the very few measurement sites.

Measurements of aerosol optical depth at several wavelengths were performed in

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the Central Mediterranean basin. These observations are used in this study combined with air mass trajectories to identify different classes of particles, and to determine their optical properties and frequency of occurrence. In a companion paper by Meloni et al. (2005) the collected observations are used to derive estimates of the aerosol single scattering albedo.

The measurements were carried out at the Station for Climate Observations, maintained by ENEA (the national Agency for New Technologies, Energy, and Environment of Italy) at Lampedusa (35.5° N, 12.6° E). Lampedusa is a small island far from continents, and the local aerosol production is limited. Thus, the aerosol properties depend essentially on long-range transport, and on the formation of marine particles in the Mediterranean. At the Station for Climatic Observations, which is located on a 45 m high plateau on the North-Eastern coast of Lampedusa, continuous observations of greenhouse gases concentration (Chamard et al., 2003), aerosol properties, total ozone, ultraviolet irradiance (di Sarra et al., 2002; Meloni et al., 2004), and other climatic parameters are carried out.

2. Instrumentation and data retrieval

This study is based on measurements obtained with a multi filter rotating shadow band radiometer (MFRSR) (Harrison et al., 1994) at Lampedusa during 2001, 2002, and 2003. The MFRSR measures global and diffuse irradiances at six 10-nm wide channels, and at one broadband channel (300–1100 nm). Measurements of global and diffuse radiation are performed every 15 s; data are averaged over longer time intervals, and stored for the analysis. The averaging interval was changed during the measurement period starting from 10 min in July 2001 to 1 min in 2003. The direct irradiance is calculated as the difference between global and diffuse irradiances, and is used to derive the aerosol optical depth in five of the six narrowband channels, centred, respectively, at 415.6, 495.7, 614.6, 672.6, and 868.7 nm. The Lambert-Beer law is applied to calculate the optical depth; the instrumental extraterrestrial constants are determined

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by means of Langley plots. The Langley method is to be used with caution at low solar elevation, because of the possible bias produced by aerosol optical depth variations. A relatively large number of extraterrestrial constant determinations are averaged to obtain reliable estimates, taking advantage of the high frequency of cloud-free conditions and, consequently, of available half-days suitable for Langley regression. The algorithm developed by Harrison and Michalsky (1994) is used to derive the instrumental extraterrestrial constant. In the considered period (July 2001–September 2003) between 65 and 169 Langley plot calibrations, depending on the channel, are averaged. First, the constants determined in the half days characterized by small values of τ are selected. They are then reduced to the mean Sun–Earth distance and averaged. Averages are calculated over the whole measurement period and over shorter time intervals to verify the instrumental stability. The constants calculated over 6-month periods differ by less than $\pm 4\%$ at 415.6 nm, and by less $\pm 2\%$ at the other channels. The instrumental characteristics remain thus reasonably constant throughout the observing period; these variations are taken into account by using yearly averages of the extraterrestrial constant. The standard deviation of the average extraterrestrial constant is always $< 2.8\%$ at 415.6 nm, and $< 2\%$ at the other channels, implying an absolute error on τ due to the uncertainty on the constant (di Sarra et al., 2002) < 0.02 . Systematic errors on τ due to erroneous values of the extraterrestrial constant are believed to remain within the estimated uncertainties. The aerosol optical depth τ at each wavelength λ is calculated by subtracting the contribution of Rayleigh scattering and ozone absorption (only for the channels at 495.7, 614.6, and 672.6 nm) from the total atmospheric optical depth. Daily mean values of atmospheric pressure and total ozone are used to derive Rayleigh and ozone optical depths (τ_{Ray} and τ_{O_3} , respectively). The ozone optical depth is calculated by using the absorption cross section at 223 K (Bogumil et al., 2003) and the total ozone amount measured routinely at Lampedusa with a Brewer spectroradiometer. When Brewer data are not available, observations from the Total Ozone Mapping Sensor, TOMS, version 8 algorithm, are used (<http://toms.gsfc.nasa.gov>). The Rayleigh optical depth is calculated following Hansen and Travis (1974) and the daily average

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pressure measured at Lampedusa. When the measured pressure is not available, a value of 1013 mbar is assumed. For this study, only cloud-free periods are selected, thus limiting the amount of available data during winter and autumn. The selection of cloud-free periods is based on the visual inspection of the irradiances (global and diffuse) measured by the broadband channel of the MFRSR. The selection method has been verified by comparing the results with visual sky observations in specific periods, and with wide angle sky photographs acquired with a total sky imager in July–August 2003.

In a companion paper (Meloni et al., 2005), radiative transfer calculations are performed at fixed values of the solar zenith angle to estimate the aerosol single scattering albedo. To maintain a uniform dataset, aerosol optical depth at fixed solar zenith angles are used also in this analysis. Averages of τ are calculated at fixed solar zenith angle, $\theta=20^\circ$, 30° , 40° , 50° , and 60° , over a $\pm 2.5^\circ$ interval around the selected θ . The $\pm 2.5^\circ$ solar zenith angle interval is chosen taking into account the measurement time resolution throughout the observation period. The uncertainty $\Delta\tau(\lambda)$ on the average value of τ is calculated as

$$\Delta\tau(\lambda) = \sqrt{(\Delta\tau_{\Delta\theta}(\lambda))^2 + (\Delta\tau_{Ray,\Delta P(Daily)}(\lambda))^2 + (\Delta\tau_{O_3(Daily)}(\lambda))^2}, \quad (1)$$

where $\Delta\tau_{\Delta\theta}$ is the standard deviation of τ in the $\pm 2.5^\circ$ solar zenith angle interval, $\Delta\tau_{Ray}$ is the daily standard deviation of τ_{Ray} resulting from the daily variability of atmospheric pressure, and $\Delta\tau_{O_3}$ is the daily standard deviation of τ_{O_3} due to changes of total ozone. When only a single measurement of τ is available in the $\pm 2.5^\circ$ solar zenith angle interval (few cases in 2001), the standard deviation of the average is substituted by 2 times the uncertainty of the single measurement of τ , calculated with the error propagation. When Brewer ozone measurements are not available, a 2% uncertainty on the TOMS total ozone is assumed. Similarly, when the local pressure measurements are missing, an uncertainty of 20 mbar (corresponding to the observed maximum annual distance from the mean value) is used. In the following analyses we discard periods characterized by a large variability of τ due to large and rapid changes of the aerosol

distribution or to residual contamination by clouds. To this aim, we remove all the data with a percent error on τ , $\Delta\tau_{\%}$, larger than the following empirical threshold $\Delta\tau_{\%M}$:

$$\Delta\tau_{\%M} = 100 * \sqrt{(0.03)^2 + (0.04 * \tau_{Ray,P_0})^2 + (0.25 * \tau_{<O3>})^2} / \tau, \quad (2)$$

where 0.03 is an empirical threshold on the variability of τ ; the other two terms on the right identify thresholds of the daily variability of Rayleigh and ozone optical depth. τ_{Ray,P_0} is the Rayleigh optical depth calculated with a pressure of 1013 mbar, and $\tau_{<O3>}$ is the ozone optical depth with a total ozone amount of 312 DU (the annual mean total ozone at Lampedusa). The behaviour of $\Delta\tau_{\%}$ and $\Delta\tau_{\%M}$ as a function of τ is shown in Fig. 1. All the data above the limiting threshold are discarded. It is worth noting that averaging around fixed values of θ implies considering different time intervals during the year; in winter long time intervals, up to 150 min, need to be used. Intervals of time shorter than 6, between 6 and 36, and longer than 36 min constitute, respectively, 10%, 84% and 6% of the dataset. The Ångström exponent, α , is the negative slope of τ versus wavelength in logarithmic scale. It is calculated from the values of the aerosol optical depth at 415.6 and 868.7 nm with the following expression:

$$\alpha = - \frac{\ln(\tau_{415.6} / \tau_{868.7})}{\ln(415.6 / 868.7)}. \quad (3)$$

The uncertainty on α is derived from the error propagation formula.

3. Analysis

3.1. Optical properties

Figure 2 shows the behaviour of α versus τ at 495.7 nm in the period July 2001–September 2003 for cloud-free conditions. Gaps in the measurements are due to the

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presence of clouds and, in minor part, to instrumental problems; the longest interruption due to an instrumental failure occurred in August 2002. A total of 1616 cloud-free measurements, corresponding to 304 days, are reported in Fig. 2. The optical depth-Ångström exponent plot, and in general the combined use of these two parameters, allows the discrimination of different aerosol types. This method has been used in a large number of studies and is based on the sensitivity of τ and α to different, somewhat independent, microphysical aerosol properties: α depends mainly on size distribution and to a lesser degree on the refractive index, while τ depends mainly on the aerosol column density. Therefore, the τ - α plot qualitatively indicates amount and size of the observed aerosol. Low and high values of α are due to the dominant presence of large and small particles, respectively, i.e. the preponderance of the coarse or the accumulation mode. The dependency of α on the size distribution has been widely used to distinguish among desert dust (e.g. Eck et al., 1999; Dubovik et al., 2002), urban-industrial (Eck et al., 1999; Smirnov et al., 2000; Dubovik et al., 2002; Kubilay et al., 2003), and biomass burning aerosols (Kaufman et al., 1998; Eck et al., 1999, 2001; Reid et al., 1999; Dubovik et al., 2002; Balis et al., 2003).

For $\tau > 0.3$, two main groups of particles can be identified in Fig. 2: the first one includes data with low values of α , typical of Saharan dust; the second one contains points with large values of α , typical of urban-industrial or biomass burning aerosols. The data point distribution of Fig. 2 is continuous, suggesting that other families of particles are present, as well as cases of mixing (i.e. coincidence) of different aerosol types.

Marine aerosols are expected to contribute significantly to the dataset as local background component. Smirnov et al. (2002) describe photometric measurements at remote oceanic sites in clear marine background conditions, and obtain values of α between 0.3 and 0.7. As discussed by Smirnov et al. (2002), at remote marine sites, the aerosol optical characteristics mostly depend on the geographical location, because of the advection of non-marine aerosol types; the influence of non-marine particles is particularly important for inland seas.

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In the following sections aerosol parameters at $\theta=60^\circ$ will be used together with air mass trajectories to identify sources and properties of particles originating from different geographical sectors.

3.2. Identification of the aerosol origin

5 Five-day back trajectories ending at Lampedusa are calculated by means of the Hysplit Dispersion Model (Draxler and Rolph, 2003) to identify the origin of the air masses; the employed model version includes vertical wind. Each trajectory is associated with the corresponding aerosol optical properties. Relating column integrated quantities, such as τ and α , to trajectories ending at specific altitudes may be problematic. Nevertheless, back trajectories are widely used in connection with aerosol optical depth (e.g. Formenti et al., 2001; Kubilay et al., 2003; Balis et al., 2003).

10 In this study we use trajectories that end over Lampedusa at 750, 2000, and 4000 m. Previous investigations based on spring-summer lidar observations (di Sarra et al., 2001) and model results (e.g. Alpert et al., 2004) have shown that African air masses are generally loaded with desert dust up to at least 4 km, and often up to 7–8 km. Conversely, when air masses originate from other regions, the aerosol is generally confined within the lowest 2–3 km of the atmosphere. The trajectories that end at 4000 m thus provide information on the transport of desert dust, that generally travels over the boundary layer; the dust load is generally large, and dominant in the atmospheric column. The trajectories ending at 2000 m are useful to identify aerosols originating from Northern sectors, while the trajectory ending at 750 m are indicative of the circulation in the lower troposphere. We define three broad geographical sectors, displayed in Fig. 3, in relation to different aerosol sources. The identified sectors are:

25 A) a southern sector, coinciding with the African continent; the dominant source is the Sahara desert;

B) a western sector, which includes the Atlantic Ocean, the Western Mediterranean Basin, the Iberian Peninsula and Western France; in this region the prevalent

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sources produce marine and, to a less extent, continental aerosols;

C) a north-northeastern sector, which includes Central and Eastern Europe; anthropogenic sources are expected to be dominant in this sector.

A degree of arbitrariness in the definition of the sectors exists, due to the effective distribution of the sources. A different identification of the sectors is thus possible. We have verified however that this simple scheme is satisfactory for the identification of broad classes of aerosol optical properties, as will be shown below.

Figure 3 shows 5-day trajectories ending at Lampedusa at 2000 m. Two trajectories per day, corresponding to $\theta=60^\circ$, are displayed for the years 2002 and 2003, when measurements of τ are available (i.e. only in cloud-free conditions, and for slowly changing aerosol optical depth, see Sect. 2). The trajectory colours in Fig. 3 indicate the value of α measured at the final point. Low values of α are measured for airmasses coming from North Africa, while airmasses arriving from sectors B and C display progressively increasing Ångström exponents. A similar dependency was found in the Eastern Mediterranean by Formenti et al. (2001), who measured aerosol properties at Mount Athos (Greece) in June-September 1998. They found that the largest values of α correspond to airmasses coming from Eastern Europe and Russia.

The aerosol origin sector can be identified by considering the time spent in the different sectors along the trajectory (di Sarra et al., 2001; Gerasopoulos et al., 2003), or by more sophisticated correlation methodologies (Formenti et al., 2001). Previous results (di Sarra et al., 2001), as well as a preliminary analyses of this dataset, indicate that dust is loaded in the air column even after a marginal passage over the African continent. Therefore, the time spent in each sector does not always provide information on the aerosol origin.

For this reason, we developed a different method to identify the aerosol source sector. We assume that the aerosol is confined to the boundary layer at the source location, and we look for regions along the trajectory where the airmass interacts with the boundary layer. We assume that aerosol is loaded when the airmass altitude, z_{air} , is

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lower or close to the altitude of the mixed layer, z_{mxl} (entrainment condition); we apply the condition $(z_{air}-z_{mxl})<500$ m. The geographical sector where this condition is met along the trajectory is identified as the source of the observed aerosol. The region around Lampedusa (1.5° latitude and longitude around the island) is excluded from the search. If the entrainment condition is met at more than one point, we choose the geographical position where the difference $(z_{air}-z_{mxl})$ is lower (sign included). Both the airmass and mixed layer altitudes are supplied by the Hysplit dispersion model.

We check first if the entrainment condition is satisfied along the trajectory ending at 4000 m; if this is not the case, the trajectories ending at 2000 m are checked. The entrainment condition for trajectories ending at 4000 m is satisfied only by African airmasses. It is interesting that all airmasses of African origin satisfy the entrainment condition for trajectories ending either at 2000 or at 4000 m.

If the entrainment condition is never met during the 5-day trajectory, we consider the time spent by the airmasses in each sector (permanence condition). The time spent in a 1.5° latitude and longitude square around the island is excluded from the calculations. We classify an airmass as originating from sectors C or B when it spent more than 50% and 75% of the last 5 days in the respective sector. The two different thresholds are determined empirically to take into account different boundary layer properties. The airmasses which do not satisfy any of those criteria are classified as of undetermined origin (D).

3.3. Back trajectories and aerosol optical properties

In this section the aerosol optical properties are related to the identified originating sectors. For this analysis we use aerosol parameters at $\theta=60^\circ$, for a total of 365 trajectories. Figure 4 shows the behaviour of α versus τ for the 365 selected data; different symbols indicate measurements classified as originating from sector A, B, C, or undetermined (D).

The distribution of the number of occurrences of α and τ is shown in Fig. 5 for each of the 4 sectors, and for different seasons. Gaussian curves are fitted to the data and

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drawn in the figure. Average values of τ and α for the different sectors, and the number and frequency of occurrence of trajectories classified in each sector, are reported in Table 1. The distributions of the values of α for sectors B and D largely overlap those for sectors A and C, which are conversely quite well separated. The aerosol optical depth progressively decreases from sector A, to C, to B/D; values belonging to the different classes largely overlap. The range of values of both τ and α for sectors B and D are very similar.

The seasonal behaviour of τ is characterized by a distinct summer increase. The value of α increases in summer for all sectors, except for sector A. The Ångström exponent is largest for sector C and smallest for sector A, indicating that the particles of sector C are smaller than those of sector A; particles of sector A are the largest. The Ångström exponent increase from winter to summer for sectors B, C, and D suggests that the optical depth increment is mainly associated with an enhancement in the small particles' amount. On the other hand, the weak seasonal change of α for sector A suggests that the summer optical depth increase is due to an enhancement in the total number of particles, and a limited increase in large particles.

About 36% of the airmasses belongs to sector A; 25% belongs to sector C, 19% to sector B, and 20% to sector D. Airmasses from sectors A and C are more frequent in summer, when the occurrence of trajectories from North-West decreases. This distribution is essentially produced by the yearly evolution of the meteorological synoptic patterns over the Central Mediterranean, with strongest North-Western flows in winter, and largest Southern influence in summer. The aerosol seasonal cycle is well explained by the synoptic patterns and by the seasonal evolution of the aerosol production mechanisms and residence time.

The desert dust and biomass burning-urban/industrial particles are well identified from the trajectory analysis, and correspond, respectively, to classes A and C. As expected, aerosols from sector A have the largest average aerosol optical depth, and the lowest Ångström exponent. The values of α of class A particles agree with those measured at Capo Verde for Saharan dust and oceanic particles (Dubovik et al., 2002). It

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is interesting that dust particles measured in the Arabic peninsula, in the Persian Gulf, or in the Hawaii display a larger range of values for α (Dubovik et al., 2002).

Sector B particles have relatively small optical depth (ranging from 0.11 to 0.19) and α between 0.5 and 1.2; these values are consistent with those of maritime particles mixed with anthropogenic aerosol or desert dust, as expected for an inland sea like the Mediterranean (Smirnov et al., 2002). As previously discussed, the ocean is considered the dominant aerosol source of the Western sector; however, this sector includes the Iberian Peninsula and Western France, and a non-homogeneous origin of the aerosol particles is expected.

The properties of aerosol from sector C correspond to urban-industrial and biomass burning particles, as identified by Dubovik et al. (2002). The largest values of α are observed during summer; as will be shown below, about 1/3 of these cases is related to forest fires occurring in Europe during the warmest and driest months of 2003. The remaining data of sector C occur in autumn-winter and spring, and display properties similar to those of sectors B and D.

It must be emphasized that different aerosol types may be present in the air column at the same time, influencing the observed optical parameters. For example, there are isolated data points belonging to sector A with a value of α as large as 1.65. As will be discussed in the following section, some of these cases are due to biomass burning products that can not be captured by a simple trajectory analysis, which does not take into account the location of fires. Similarly, there are data points classified as C with $\alpha=0.4$, or classified as B with large optical depth and low Ångström exponent.

An interesting example of mixing has been discussed by Eck et al. (1999). They show a case with values of α and τ at 500 nm, respectively, of 1 and 0.43, resulting from the mixing of desert and industrial aerosols; the mixing led to a size distribution having accumulation and coarse mode of similar amplitude.

We examined all the trajectories that produce outliers in the data distribution of each specific sector to verify the coherence between the classification based on the airmass trajectories and the aerosol optical parameters. Most of the outliers correspond to rela-

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tively fast changes of the airmass path occurring within about one day. These changes determine a weakening of the relation between the trajectory and the optical characteristics; a large uncertainty on the trajectory path or on the estimate of the mixing layer altitude, a large variability of the aerosol properties, and/or a large inhomogeneity of the aerosol vertical distribution, are among the possible causes.

In the following section we discuss the summer 2003 observations, when unusual meteorological conditions influenced the aerosol load over the Central Mediterranean basin. In this period desert dust and biomass burning particles are clearly identifiable. Mixed aerosol types are also observed; they correspond to many of the outliers of Fig. 4.

3.4. Case study: July-August 2003

Figure 6 shows the mean values of τ and α at different solar zenith angles during the period June 28 (day 179)–September 1 (day 244), 2003. Different symbols indicate the identified airmass origin at $\theta=60^\circ$. Three intervals with $\alpha \geq 1.5$ (periods I, II, and III) are highlighted in the figure. Values of τ are generally >0.2 during July and August, and are ≥ 0.4 around days 181, 196–198, 204–205, 218–225, and 243–245; α varies from -0.2 to $+2$. The longest period with persistent large values of α lasts from day 212 to 232, when τ is generally >0.3 ; these data belong to class C, and are unusual in the whole dataset. It is thus worth investigating their possible causes.

Summer 2003 has been the European hottest summer since 1500 (Luterbacher et al., 2004). A persistent high pressure system remained centred over Western Europe during most of July and August, contributing to maintain high temperatures and dry weather. These conditions were favourable to the enhancement in the atmospheric aerosol concentration, and to the development of forest fires: huge, ravaging and persistent fires developed in Portugal, Provence, Italy, and the Balkan region; such fires and their plumes are well documented in satellite images (see e.g. <http://visibleearth.nasa.gov>). The exceptionally high temperature, as well as the increase in the ambient ozone and aerosol concentration led to a number of deaths

higher than normal (e.g. Fisher et al., 2004; Stedman, 2004).

Large amounts of biomass burning particles were driven from the European continent to the Central-Southern Mediterranean by the high pressure system. Eck et al. (1999, 2001) have shown that biomass burning aerosol is characterized by large values of α , generally >1.5 . Biomass burning aerosol in Brazil, Zambia, and Canada were observed to produce mean values of τ at 440 nm between 0.4 and 0.8, and values of α in the range from 1.2 to 2.2 (Dubovik et al., 2002). Recently Balis et al. (2003) reported observations of smoke plumes originating from different European regions (Southern Italy, Danube region and Northern Black Sea). They measured instantaneous values of τ as large as 0.84 at 440 nm, and values of α between 1.78 and 1.82 in August 2001 at Thessaloniki (Greece). Urban/industrial aerosols display large values of α : generally $\alpha > 1.4$ in Southern Europe (Holben et al., 2001; Gerasopoulos et al., 2003).

To identify sources of biomass burning particles we use the Web Fire Mapping that provides daily maps of fires (<http://maps.geog.umd.edu/default.asp>, a collaboration between NASA Goddard Space Flight Center and University of Maryland based on the MODerate resolution Imaging Spectroradiometer, MODIS, data). The daily maps are combined with the airmass trajectories. Figure 7 shows 5-day trajectories ending at Lampedusa at 2000 m in the period 28 June–1 September 2003. The colour code indicates the value of α .

Observations of the period I are discussed in the companion paper by Meloni et al. (2005), and we focus here on periods II and III. In period II the trajectories at all altitude levels overpass Sicily and Southern Italy, where several fires were active during the previous days. In period III almost all airmasses flowed over Sicily, where fires were active, before approaching Lampedusa. During the previous days trajectories overpassed different regions with forest fires: the East Adriatic coast, South-Eastern Italy, Sardinia, and the Northern coast of Africa. Thus, we attribute the values of $\alpha \geq 1.5$ to biomass burning particles; urban sources may constitute an additional, although not dominant, contribution. To our knowledge, fire plume episodes of such a large temporal

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and spatial extension have not been previously reported in the Mediterranean.

Meloni et al. (2005) show that the retrieved aerosol single scattering albedo of periods II and III corresponds to biomass burning, coherently with our finding. Their analysis also shows that biomass burning or urban/industrial particles are dominant in the aerosol column on different days.

Before and after each of the three periods, Saharan dust outbreaks lead to rapid changes of the aerosol characteristics. In the transition periods, particles displaying intermediate/mixed optical properties appear. In particular, at the beginning of period III (days 208–210) airmasses are classified as A; the observed value of α is between 0.5 and 0.8, and τ is about 0.3. Such a large value of τ is usually connected to dust or heavily polluted/fire particles; however, the corresponding value of α does not allow to discriminate between the two aerosol types. MODIS satellite images show a large desert haze over the sea, West and North-West of Sardinia and Corsica, and a large smoke plume from fires over Central Corsica on 22 July (day 203). According to the trajectories, the airmass arriving at Lampedusa at 2000 m on 27 July (day 208), loaded with desert dust, captured smoke particles overpassing Corsica on 22 July. Thus, the observed optical parameters (daily average $\tau=0.28$ and $\alpha=0.81$) are the result of mixing of different aerosol types. Similar cases are also documented by satellite imagery. The Web Fire Mapping also shows frequent fires along the Northern coast of Africa in August 2003; their presence probably explains the cases of particles classified as A displaying $\alpha\sim 0.75$ and $\tau>0.2$.

4. Optical properties and aerosol types

In this section the observed values of τ and α (mean values at fixed solar zenith angles, $\theta=20^\circ$, 30° , 40° , 50° , 60°) are studied with respect to the known optical properties of desert dust and biomass burning-urban/industrial aerosol types, hereafter indicated as DD and BU, respectively. These two classes of particles produce significantly different radiative effects: in the visible spectral range scattering is prevalent for DD, and absorp-

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tion for BU particles. Thus, the identification of cases with a single dominant aerosol type is essential for the assessment of the aerosol radiative effects in the Mediterranean. The companion paper by Meloni et al. (2005) is dedicated to the determination of the aerosol single scattering albedo; their analysis is based on the following classification of the main aerosol types.

There is rarely a single aerosol type in the air column over the Mediterranean, since marine particles are usually dominant in the boundary layer, and other types mix and/or travel above. However, we identify cases dominated by a single type when the values of τ and α are within specific conservative ranges; these ranges are defined using typical values for desert dust and biomass burning-urban/industrial aerosol derived from the literature (see the discussion in Sect. 3.3), and the distribution of the observed values of τ and α at Lampedusa (see Fig. 4). Cases dominated by DD are those with $\tau \geq 0.15$ and $\alpha \leq 0.5$; cases dominated by BU particles are those with $\tau \geq 0.1$ and $\alpha \geq 1.5$. Although other types might be identified, it would not be possible to distinguish cases dominated by a single type from cases of mixed aerosol. Thus, we define only two aerosol types, and classify the remaining cases as mixed (M).

It is worth emphasizing that about 90% of the class DD data belongs to sector A, 6% to sector B, and only 1 and 3% to sector C and class D, respectively. Conversely, 65% of the BU data originates from sector C, 13% from sector B and 2% from sector A, while 20% belongs to class D. As expected, DD particles have a more defined geographical origin, and frequently produce conditions dominated by a single aerosol type. The reader is reminded that these statistics are derived from trajectories ending at Lampedusa at $\theta = 60^\circ$.

Figures 8 and 9 display the number of occurrences of α and τ for DD, BU, and M aerosol types for the whole dataset and in different seasons. Three summer data points with $\tau > 0.9$ are excluded from Figs. 8 and 9 to increase the readability of the histograms. Due to the annual variability of cloudiness, the amount of available data is not uniform throughout the year, and the statistical significance of the results is larger in spring and summer. Autumn and winter data together constitute only 6.1% of the dataset, and are

grouped together. Table 2 reports number and frequency of occurrence, and mean and standard deviation of the distribution of τ and α .

The overall mean and standard deviation of τ and α are 0.24 ± 0.14 and 0.86 ± 0.63 , respectively. The values of τ range from 0.03 to 1.13, while the values of α vary from -0.32 to 2.05. The average values of τ and α are influenced by the uneven temporal distribution of the data, and their significance must be considered with caution. For this reason we calculated three different averages, whose results are reported in Table 3: the first is calculated over the whole dataset; the second is calculated over two entire annual cycles; the third is calculated from daily averages, using all the available cloud-free periods, and over two entire annual cycles. The third average value is smaller than the other two, that contain a larger number of observations at solar zenith angles $< 60^\circ$ during spring and summer.

Gerasopoulos et al. (2003) discuss observations of aerosol properties at a rural (Ouranoupolis 40.4° N, 23.9° E) and an urban site (Thessaloniki, 40.6° N, 22.9° E) in Northern Greece. They report averages of aerosol optical depth at 500 nm: $\tau = 0.23 \pm 0.13$ at the rural site, and 0.23 ± 0.15 at the urban site; their results are very close to those obtained in this study. It has to be noted that the averages are calculated by Gerasopoulos et al. (2003) over the entire datasets, and may differ from annual means. Average values of α are much larger (1.4 ± 0.4 and 1.6 ± 0.3 at the rural and urban site, respectively) than at Lampedusa. This difference is attributed to the different geographical locations and the corresponding prevalent aerosol types. Northern Greece is strongly influenced by regional (anthropic) sources, and occasionally experiences cases of long range transport of desert dust. Conversely, the geographic position of Lampedusa determines a dominant role of long-range transport, and, due to the proximity of Africa, a stronger influence from desert dust.

As it is evident in Fig. 9, the optical depth seasonal evolution of all aerosol types is characterized by a summer increase. The presence of aerosol classified as BU is a distinct summer phenomenon, since very few cases occur in autumn-winter, and no cases in spring. Desert dust is known to display a clear seasonal cycle in the Central

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Mediterranean (Moulin et al., 1998; Israelevich et al., 2002), with a distinct summer maximum. The increase in the optical depth of the other aerosol types is related to reduced precipitation, stronger convection, and longer permanence of particles in the atmosphere during summer. Gröbner and Meleti (2004) report observations of aerosol optical depth at 443 nm at Ispra (Italy) during the period 1995–2002; their data show a distinct annual cycle, with a maximum of 0.5 in June, and a minimum of 0.18 in November–December. Multiyear measurements with MFRSR radiometers in Central and Eastern US show a summer increase in τ and α , followed by a reduction in autumn-winter; the annual average α is close to 1.3, representative of continental aerosols (Michalsky et al., 2001). They show that the seasonal cycle is larger in Eastern coastal regions than the rest of the US. The seasonal behaviour in the US is confirmed by other climatological studies (Holben et al., 2001; Alexandrov et al., 2002). Holben et al. (2001) extend this result to all Northern midlatitude continental sites.

The summer increase in τ and α at Lampedusa is consistent with other Sun photometric measurements at Southern Europe sites. Observations at Ispra (Italy, 45.8° N, 8.6° E) during 1997–1998 show a monthly mean maximum of $\tau=0.52$ at 500 nm in June, and two maxima of α (calculated between 440 and 870 nm) in May and July (1.62 ± 0.15 and 1.61 ± 0.13) (Holben et al., 2001). Monthly mean values of α measured at Tito Scalo (Italy, 40.6° N, 15.7° E) decrease from 1.73 ± 0.25 in June, to 0.90 ± 0.03 in October 2001, confirming the reduction in the content of small particles from summer to fall (Esposito et al., 2004). Gerasopoulos et al. (2003) show that maxima of τ and α monthly averages occur in August (~ 0.3 and ~ 1.6 , respectively), and minima in December (~ 0.09 and ~ 1.0 , respectively) at a rural site in Northern Greece. They find a different behaviour at the Greek urban site, where no clear annual cycle appears; monthly values of α are generally large (ranging from ~ 1.3 to ~ 1.7), possibly due to the influence of local pollution.

We can not exclude that part of the seasonality observed at Lampedusa is induced by the temporal distribution of cloudiness. However, the behaviour of the urban-industrial and biomass burning aerosol is essentially determined by the accumulation mode,

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and observations reflect the longer residence time, the stronger efficiency of sources (mainly for forest fires), and the stronger convective regime occurring in summer. Entrainment of aerosol particles in the free troposphere, allowing for long range transport, may also play a role. Entrainment in the free troposphere is stronger in the convective regimes characteristic of the summer season in Central-Southern Europe, while it is normally reduced in the temperature inversion regime typical of colder months.

The increased capability of the summer advection regime to transport aerosols of European origin to the Mediterranean is discussed by Stohl et al. (2002) and Duncan and Bey (2004), who show that the main outflow toward the Mediterranean Basin and Africa occurs during summer; this flow is strongly reduced during the other seasons.

Multi-year aerosol optical depth estimates were derived from satellite observations. The retrieval of the optical depth from satellite data depends on some assumptions. Due to these assumptions and to the interannual variability of the aerosol distribution (e.g. Moulin et al., 1997a), a direct comparison is problematic. However, satellite data are available over extended time periods, and a comparison of our results with these climatological values may provide useful indications.

Using 11 years of Meteosat data Moulin et al. (1998) estimate the African dust seasonal contribution to the aerosol optical depth over the Central Mediterranean; they retrieve a seasonal average dust optical depth at 550 nm of 0.08 ± 0.02 , 0.15 ± 0.06 , 0.18 ± 0.06 , 0.09 ± 0.04 , respectively, for winter, spring, summer, and autumn; the annual mean is 0.13 ± 0.04 . The aerosol optical depth at 550 nm is calculated from the MFRSR observations at Lampedusa using the Ångström law. Averaging daily values, we obtain an annual mean of 0.20 ± 0.13 , and seasonal averages of 0.15 ± 0.11 , 0.16 ± 0.11 and 0.25 ± 0.13 for autumn-winter, spring and summer, respectively. Our dataset refers to a different and shorter time period, and the comparison should be considered with caution. However, we measure substantially larger values of τ , and a less pronounced dust seasonal cycle. These differences are partially explained by Moulin et al. (1997b), who show that the Meteosat retrieval underestimates the aerosol optical depth determined from Sun photometers.

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Husar et al. (1997) derived monthly average aerosol optical depths from AVHRR (Advanced Very High Resolution Radiometer) observations at 630 nm for the period July 1989–June 1991. They obtain an annual mean of 0.18 in the Mediterranean basin; the annual aerosol optical depth cycle has a minimum of about 0.07 in December, and a maximum of about 0.28 in August. Averaging their monthly aerosol optical depths we derive values of about 0.11, 0.23, and 0.26 for autumn-winter, spring and summer, respectively. Our measurements of τ at 615 nm give an annual mean of 0.19 ± 0.13 , and values of 0.14 ± 0.11 , 0.15 ± 0.11 and 0.23 ± 0.13 for autumn-winter, spring and summer, respectively. The annual averages are very similar, but we observe a less pronounced annual cycle, with larger values of τ during autumn-winter, and lower during spring and summer. It must be pointed out that the results by Husar et al. (1997) are integrated over the whole basin, and significant differences in the aerosol seasonal behaviour in the various parts of the Mediterranean exist (see e.g. Moulin et al., 1998; Israelevich et al., 2002).

5. Conclusions

At the island of Lampedusa ENEA maintains a Station for Climate Observations, where measurements of greenhouse gases, aerosol properties, total ozone, ultraviolet and solar radiation and other climatic parameters are carried out. Lampedusa is a small island located in the Central Mediterranean, East of Tunisia and South of Sicily, without significant local aerosol sources. In this study we present almost 2.5 years of aerosol optical depth and Ångström exponent measurements obtained with a multi filter rotating shadowband radiometer (MFRSR).

The main results of this analysis may be summarized as follows:

1. The average aerosol optical depth at 495.7 nm and the average Ångström exponent are 0.24 ± 0.14 and 0.86 ± 0.63 , respectively, over the period July 2001–September 2003. The observed values of τ range from 0.03 to 1.13, and the

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values of α vary from -0.32 to 2.05 , indicating a large variability in aerosol content and size.

2. The airmasses were classified according to the originating sectors during the cloud-free periods. 36% of the airmasses come from Africa, 25% from Central-Eastern Europe, and 19% from Western France, Spain and the North Atlantic. In summer, 42% of the airmasses is of African origin. In almost all cases African aerosols display high values of τ and low values of α , typical of Saharan dust (average values of τ and α are 0.36 and 0.42, respectively). Particles originating from Central-Eastern Europe show relatively large average values of τ and α (0.23 and 1.5, respectively), while particles from Western France, Spain and the North Atlantic show the lowest average values of τ (0.15), and relatively small values of α (0.92).
3. Cases of intermediate values of α are often related to relatively fast changes in the airmass originating sector, suggesting the contemporary presence of different types of particles in the air column. The largest values of α (about 2) are observed in August 2003, when intense forest fires in Southern Europe produced large amounts of fine combustion particles, that were transported to the Central Mediterranean by a persistent high pressure system over Central Europe. Smoke particles in some cases mix with desert dust, producing intermediate values of α .
4. The seasonal distribution of the meteorological patterns over the Mediterranean, the efficiency of the aerosol production mechanisms, and the variability of the particles' residence time in the atmosphere produce a distinct seasonal cycle of aerosol optical depths and Ångström exponent values. Particles originating from all sectors show a summer maximum in aerosol optical depth. The summer increase in optical depth for European aerosols is linked with an increment in the values of α , that indicates an enhancement in the number of fine particles. The summer maximum of τ for African particles is associated with a weak reduction

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in the Ångström exponent, suggesting an increase in the total number of particles and a relatively more intense transport of large particles.

- 5 The observations were classified according to the aerosol optical properties, and two main classes have been identified: desert dust (DD) and biomass burning/urban-industrial (BU) aerosols. Values of τ and α averaged over the whole observing period are 0.37 and 0.15 for DD, and 0.27 and 1.77 for BU. 90% of the cases classified as DD originate from Africa, while 65% of the BU cases originate from Central-Eastern Europe. A seasonal evolution of the aerosol properties of the different types is also evident.
- 10 6. The comparison with observations at other Mediterranean sites (mostly located on the coasts of Europe) reveals that desert dust plays an important role at Lampedusa.

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Table 1. Mean, standard deviation and number (frequency) of occurrences of τ and α , for airmasses originating from different sectors. Data are averaged around the solar zenith angle of 60° .

| Sector | Whole period | Autumn-winter | Spring | Summer |
|--------------------------------------|----------------|----------------|----------------|----------------|
| Number of occurrences | 365 | 42 (11.5%) | 90 (24.6%) | 233 (63.8%) |
| A (Northern Africa) | | | | |
| Number of occurrences | 132 (36%) | 10 (24%) | 24 (27%) | 98 (42%) |
| τ | 0.36 ± 0.16 | 0.25 ± 0.14 | 0.30 ± 0.16 | 0.38 ± 0.16 |
| α | 0.42 ± 0.40 | 0.50 ± 0.46 | 0.39 ± 0.38 | 0.42 ± 0.40 |
| B (Atlantic – Western Europe) | | | | |
| Number of occurrences | 69 (19%) | 15 (35%) | 24 (27%) | 30 (13%) |
| τ | 0.15 ± 0.08 | 0.15 ± 0.10 | 0.12 ± 0.04 | 0.18 ± 0.07 |
| α | 0.92 ± 0.44 | 0.62 ± 0.38 | 0.80 ± 0.29 | 1.16 ± 0.43 |
| C (Central – Eastern Europe) | | | | |
| Number of occurrences | 92 (25%) | 4 (10%) | 14 (15%) | 74 (31%) |
| τ | 0.23 ± 0.11 | 0.11 ± 0.02 | 0.19 ± 0.06 | 0.25 ± 0.12 |
| α | 1.50 ± 0.36 | 0.95 ± 0.09 | 1.06 ± 0.14 | 1.60 ± 0.30 |
| D (Undetermined) | | | | |
| Number of occurrences | 72 (20%) | 13 (31%) | 28 (31%) | 31 (13%) |
| τ | 0.14 ± 0.06 | 0.12 ± 0.08 | 0.12 ± 0.05 | 0.20 ± 0.06 |
| α | 1.02 ± 0.52 | 0.51 ± 0.34 | 0.85 ± 0.38 | 1.39 ± 0.43 |

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Table 2. Mean, standard deviation and number (frequency) of occurrences of τ and α , for the identified aerosol types. Data are averaged around 20°, 30°, 40°, 50°, and 60° solar zenith angle.

| Aerosol type | Whole period | Autumn-winter | Spring | Summer |
|---|--------------|---------------|-----------|------------|
| Whole dataset | | | | |
| Number of occurrences | 1616 | 99 (6%) | 373 (23%) | 1144 (71%) |
| τ | 0.24±0.14 | 0.18±0.12 | 0.18±0.18 | 0.27±0.14 |
| α | 0.86±0.63 | 0.61±0.41 | 0.67±0.42 | 0.95±0.68 |
| DD (Desert dust) | | | | |
| Number of occurrences | 475 (30 %) | 26 (26%) | 83 (23%) | 366 (32%) |
| τ | 0.37±0.14 | 0.30±0.12 | 0.33±0.14 | 0.39±0.14 |
| α | 0.15±0.17 | 0.22±0.16 | 0.13±0.17 | 0.16±0.17 |
| BU (Biomass burning/ urban-industrial) | | | | |
| Number of occurrences | 309 (19%) | 2 (2%) | 1 (0.03%) | 306 (27%) |
| τ | 0.27±0.11 | 0.13±0.01 | 0.31±0.01 | 0.27±0.11 |
| α | 1.77±0.15 | 1.69±0.03 | 1.55±0.05 | 1.78±0.15 |
| M (Mixed) | | | | |
| Number of occurrences | 832 (51%) | 71 (72%) | 289 (77%) | 472 (41%) |
| τ | 0.16±0.09 | 0.13±0.09 | 0.13±0.06 | 0.19±0.10 |
| α | 0.93±0.37 | 0.72±0.36 | 0.82±0.32 | 1.02±0.34 |

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Table 3. Mean and standard deviation of τ and α calculated with different data selection criteria.

| Period | Jul. 2001–Sep. 2003 | Aug. 2001–Aug. 2003 | Aug. 2001–Aug. 2003 |
|--------------|---------------------|---------------------|---------------------|
| Type of data | Whole dataset | Whole dataset | Daily means |
| τ | 0.24±0.14 | 0.23±0.14 | 0.22±0.13 |
| α | 0.86±0.63 | 0.87±0.64 | 0.82±0.59 |

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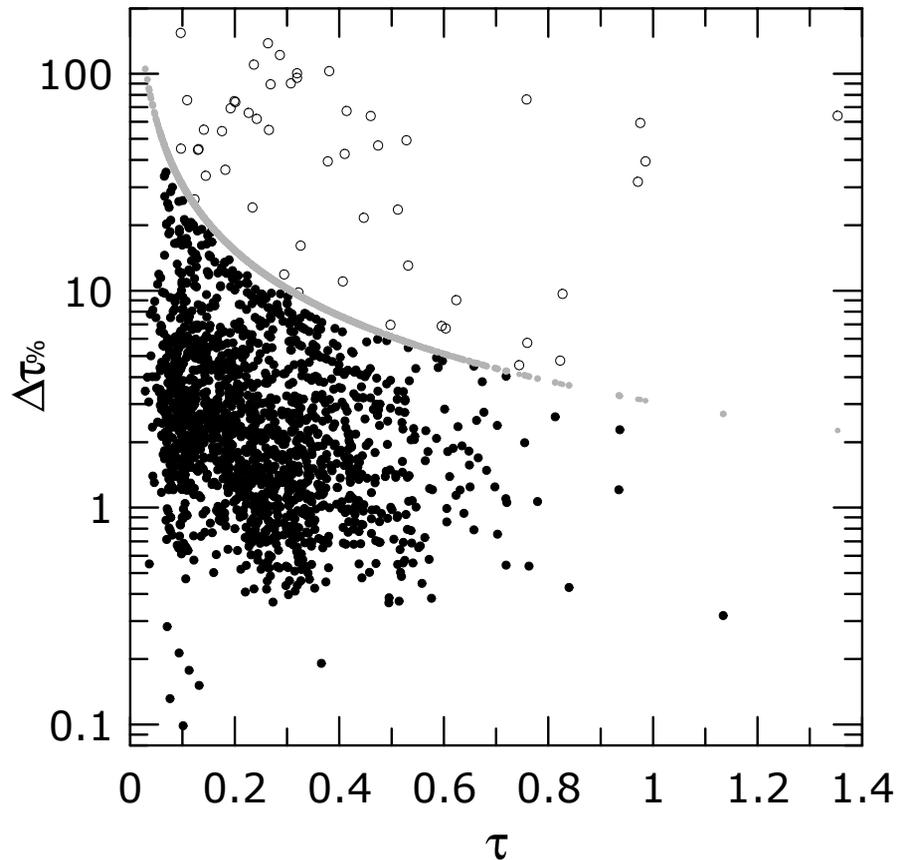


Fig. 1. Behaviour of the percent error, $\Delta\tau\%$, versus the aerosol optical depth at 495.7 nm, τ . Grey dots are the empirical threshold used to discriminate residual points with large variability; full and open black circles are the selected and excluded measurements, respectively.

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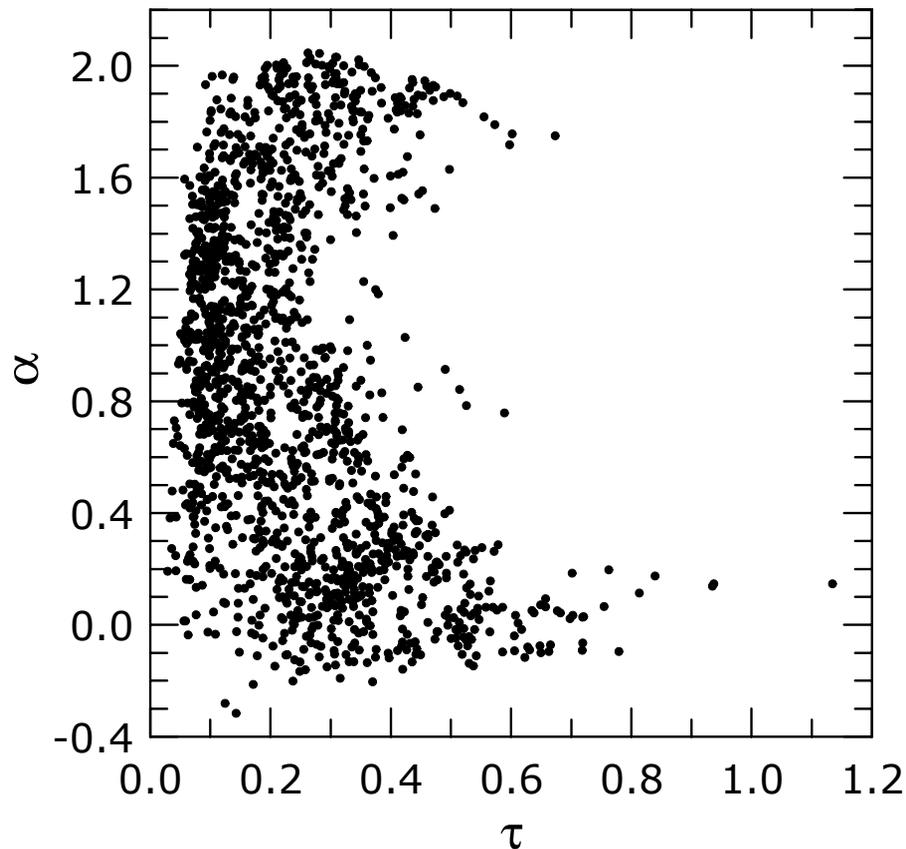


Fig. 2. Behaviour of the Ångström exponent, α , as a function of the aerosol optical depth at 495.7 nm, τ , for cloud-free conditions during July 2001–September 2003.

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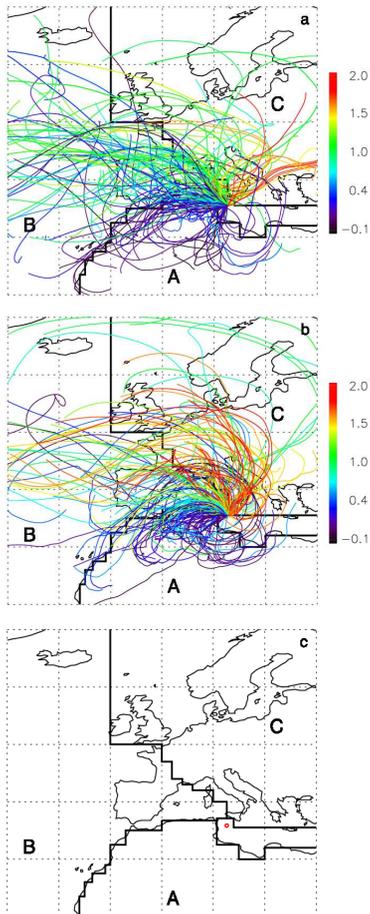


Fig. 3. Five-day trajectories ending at Lampedusa at 2000 m, in **(a)** 2002, and **(b)** 2003. The colour of the trajectories refers to the corresponding value of α measured at Lampedusa. Two trajectories per day, corresponding to a solar zenith angle of 60° , are displayed. The different source regions are split into three sectors (A, B, and C) shown in panel **(c)**. The position of Lampedusa is indicated by a red circle.

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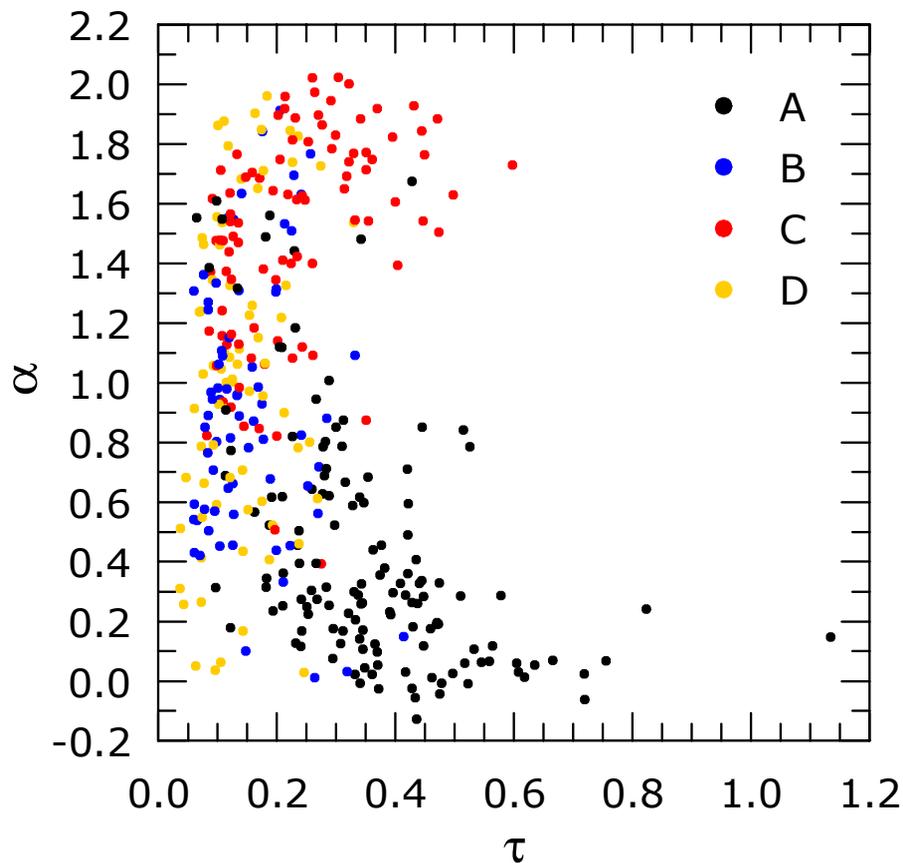


Fig. 4. Behaviour of the Ångström exponent, α , as a function of the aerosol optical depth at 495.7 nm, τ . Only data corresponding to a solar zenith angle of 60° are considered. The colour of the symbols indicates the airmass originating sector.

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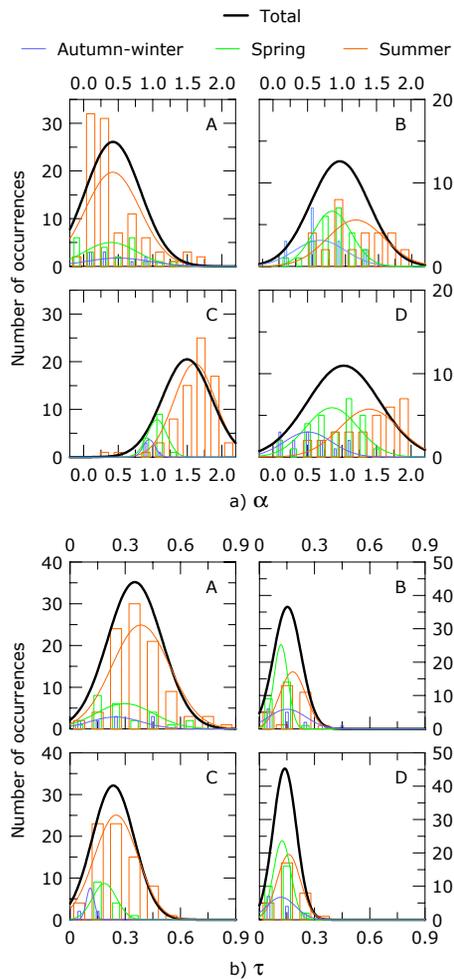


Fig. 5. Distribution of the number of occurrences of (a) α , and (b) τ , for each of the 4 sectors: A, B, C, D. The seasonal data are shown in orange (Summer), green (Spring) and blue (Autumn-Winter). Distributions for the different seasons (coloured lines) and for the whole dataset (black thick line) are identified by Gaussian curves fitted to the data.

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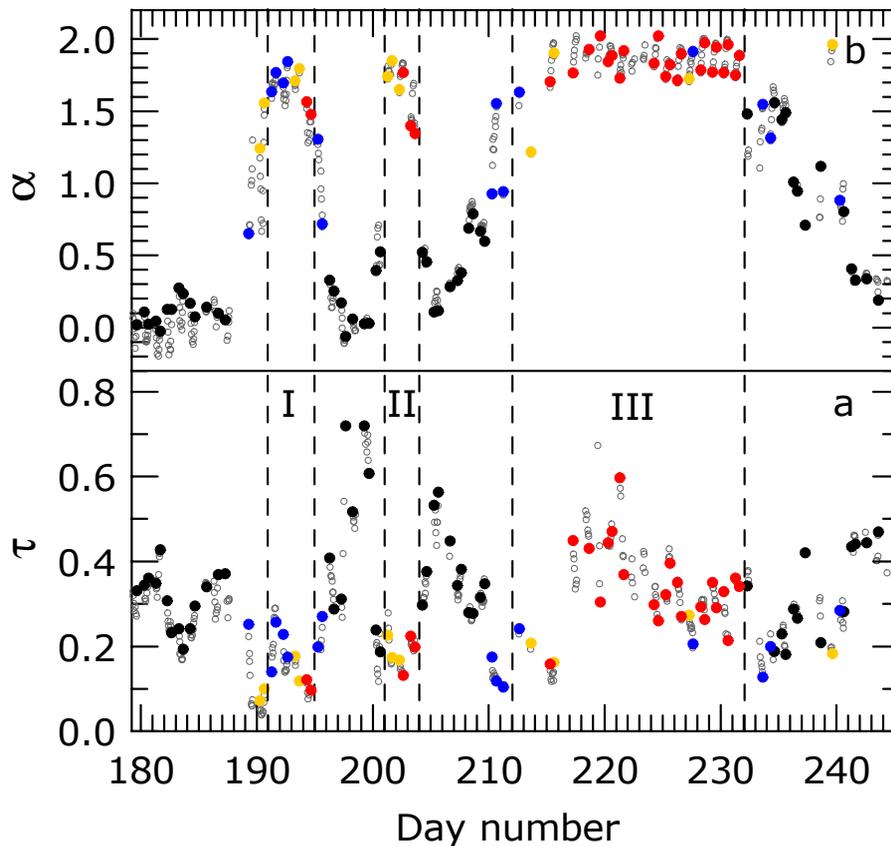


Fig. 6. Behaviour of (a) τ at 495.7 nm, and (b) α (values at 60° solar zenith angle are indicated by solid circles in black, blue, red, and yellow; open grey circles indicate values at other solar zenith angles) in the period 28 June–1 September 2003. The colours of the solid circles identify the origin of the air mass according to the trajectory analysis (see Fig. 4).

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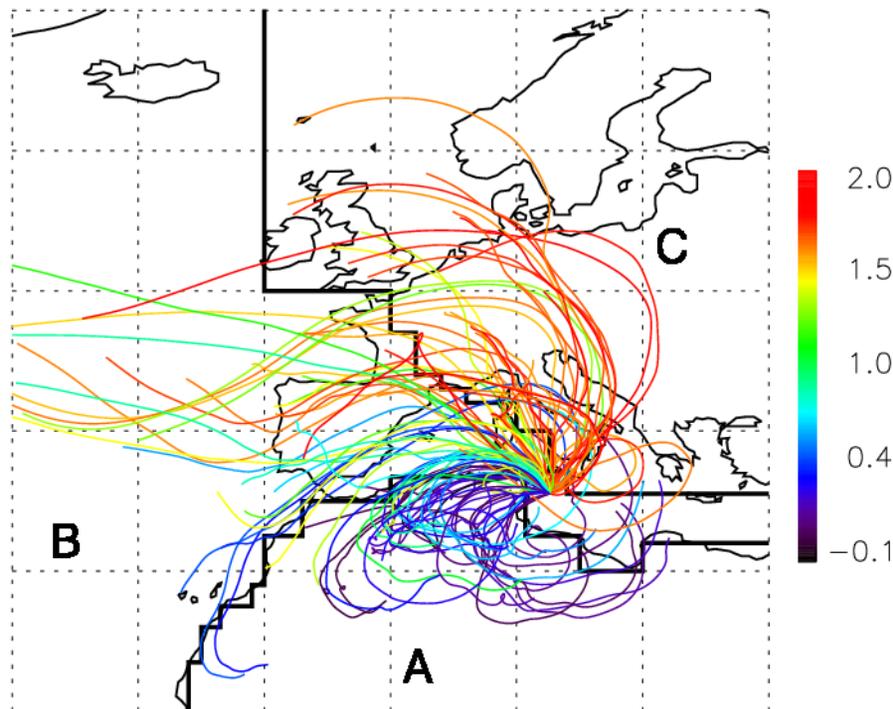


Fig. 7. Five-day trajectories ending at Lampedusa at 2000 m in the period 28 June–1 September 2003. Two trajectories per day corresponding to values of the solar zenith angle of 60° are displayed, except in cloudy periods. The colour of the curves indicates the value of α measured at Lampedusa at the end of the trajectory.

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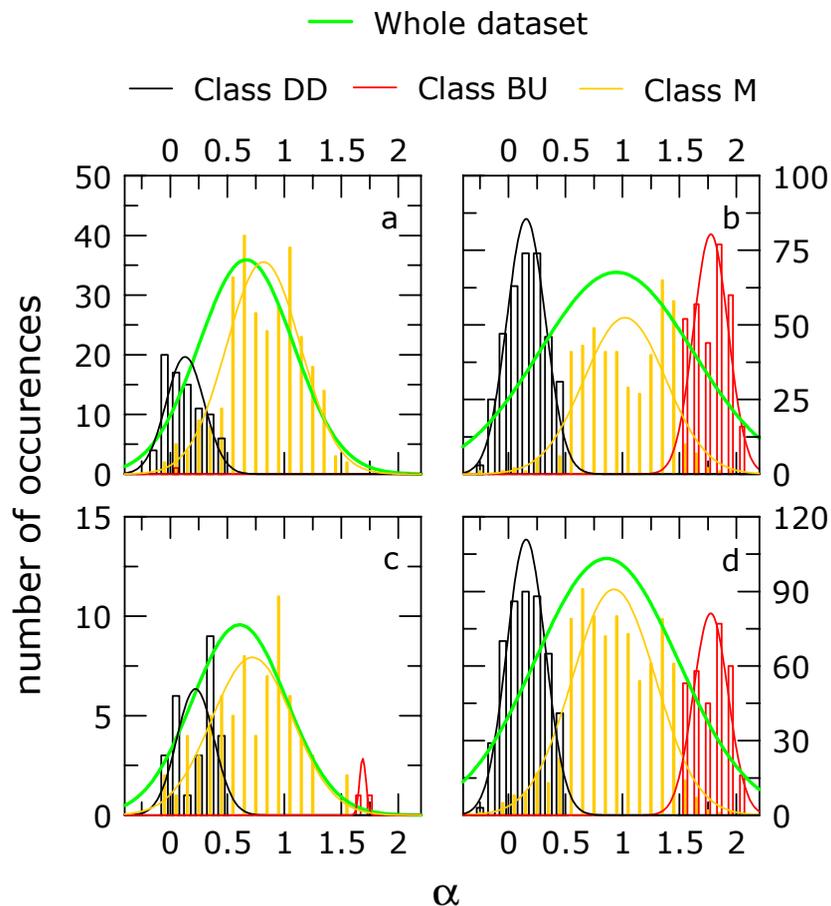
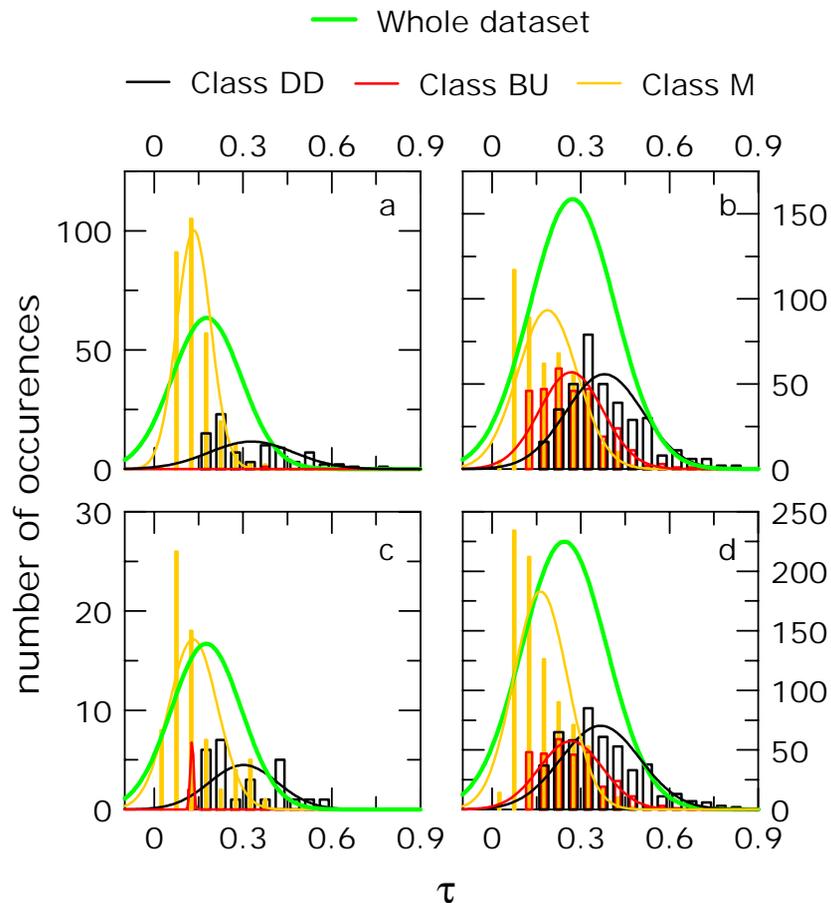


Fig. 8. Spring (a), summer (b), autumn-winter (c), and the whole dataset (d), distribution of the number of occurrences of α . Histograms and Gaussian curves fitted to the data for the DD, BU and M classes are indicated by black, red, and yellow lines, respectively. Green tick lines indicate Gaussian curves fitted to the whole dataset.

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**Fig. 9.** Same as Fig. 8, but for τ .[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)