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# Divergence of the Ensemble Transform Kalman Filter (LETKF) by nonlocal observations

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## 2 ABSTRACT

3 Ensemble Kalman filters are powerful tools to merge model dynamics and observation data. For  
4 large system models, they are known to diverge due to subsampling errors at small ensemble  
5 size and thus possible spurious correlations in forecast error covariances. The Local Ensemble  
6 Transform Kalman filter (LETKF) remedies these disadvantages by localisation in observation  
7 space. However, its application to nonlocal observations is still under debate since it is still not  
8 clear how to optimally localize nonlocal observations. The present work studies intermittent  
9 divergence of filter innovations and shows that it increases forecast errors. Nonlocal observations  
10 enhance such innovation divergence under certain conditions, whereas similar localisation radius  
11 and sensitivity function width of nonlocal observations minimizes the divergence rate. The analysis  
12 of the LETKF reveals inconsistencies in the assimilation of observed and unobserved model  
13 grid points which may yield detrimental effects. These inconsistencies *inter alia* indicate that the  
14 localisation radius should be larger than the sensitivity function width if spatially synchronised  
15 system activity is expected. Moreover, the shift of observation power from observed to unobserved  
16 grid points hypothesised in the context of catastrophic filter divergence is supported for intermittent  
17 innovation divergence. Further possible mechanisms yielding such innovation divergence are  
18 ensemble member alignment and a novel covariation between background perturbations in  
19 location and observation space.

20 **Keywords:** ensemble Kalman filter, localisation, nonlocal observations, divergence, local observations

## 1 INTRODUCTION

21 Data assimilation (DA) merges models and observations to gain optimal model state estimates. It is well-  
22 established in meteorology [1], geophysics [2] and attracts attention in life sciences [3]. Typical applications  
23 of DA serve to estimate model parameters [4] or provide initial conditions for forecasts [5]. A prominent  
24 technique is the ensemble Kalman filter [6], which allows to assimilate observations in nonlinear models.  
25 When underlying models are high-dimensional, such as in geophysics or meteorology, spurious correlations  
26 in forecast errors are detrimental to state estimates. A prominent approach to avoid this effect is localisation  
27 of error covariances. The Local Ensemble Transform Kalman Filter (LETKF) [7] utilises a localisation  
28 scheme in observation space that is computationally effective and applicable to high-dimensional model

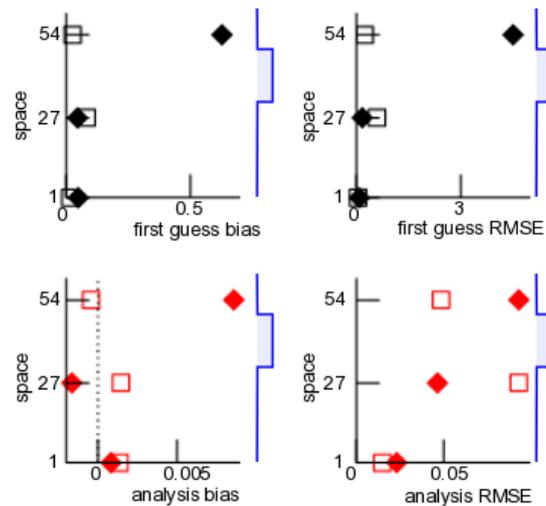
29 systems. The LETKF applies to local observations [8] measured in the physical system under study, e.g. by  
30 radiosondes, and nonlocal observations measured over a large area of the system by, e.g., weather radar or  
31 satellites [9, 10, 11]. Since nonlocal observations represent spatial integrals of activity, and the localisation  
32 scheme of the LETKF requests a single spatial location of each observation, it is conceptually difficult to  
33 apply the LETKF to nonlocal observations. In fact, present localisation definitions [12, 10] of nonlocal  
34 observations attempt to estimate the best single spatial location neglecting the spatial distribution of possible  
35 activity sources. A recent study [13] on satellite data assimilation proposes to choose the localisation radius  
36 equal to the spatial distribution width of radiation sources. This spatial source distribution is the sensitivity  
37 function of the nonlocal observation and is part of the model system. The present work considers the  
38 hypothesis that the relation between localisation radius and sensitivity function width plays an important  
39 role in the filter performance.

40 Merging the model forecast state and observations, the ensemble Kalman filter tears the analysis, i.e.  
41 the newly estimated state, towards the model forecast state and thus underestimates the forecast error  
42 covariance matrix due to a limited ensemble size [14]. This is enforced by model errors [15, 16] and leads  
43 to filter divergence. Moreover, if the forecast error covariances are too large, the forecasts have too less  
44 weight in the assimilation step and the filter tears the analysis towards the observations. This also results to  
45 filter divergence. In general terms, filter divergence occurs when an incorrect background state can not be  
46 adjusted to a better estimate of the true state by assimilating observations.

47 Ensemble member inflation and localisation improves the filter performance. The present work considers a  
48 perfect model and thus neglects model errors. By virtue of this study construction, all divergence effects  
49 observed result from undersampling and localisation. The present work chooses a small ensemble size  
50 compared to the model dimension, fixes the ensemble inflation to a flow-independent additive inflation and  
51 investigates the effect of localisation.

52 In addition to the filter divergence described above ensemble Kalman filter may exhibit catastrophic filter  
53 divergence which enhances the filter forecasts to numerical machine infinity [17, 18, 19, 20, 21]. This  
54 divergence is supposed to result from alignment of ensemble members and from unconserved observable  
55 energy dissipation [20]. This last criterion states that the filter diverges in a catastrophic manner if the  
56 observable energy of the system dissipates in unobserved directions, i.e. that energy moves from observed  
57 to unobserved locations. The present work raises the question whether such features of catastrophic  
58 divergence play a role in non-catastrophic filter divergence as well. Subsequent sections indicate that this is  
59 the case in the assimilation of nonlocal observations.

60 The underlying motivation of this work is the experience from meteorological data assimilation, that  
61 satellite data are detrimental to forecasts if assimilation procedure is not well-tuned [12, 13, 22]. This effect  
62 is supposed to result from deficits in the underlying model. The present work assumes a perfect model and  
63 investigates the question, whether assimilating nonlocal observations is still detrimental. Figure 1 shows  
64 forecast and analysis errors in numerical data assimilation experiments with this perfect model with three  
65 local observations only and with additional nonlocal observation. Nonlocal observations have positive and  
66 negative impact on the forecast error of the local observations dependent on the spatial location of the local  
67 observations with respect to the nonlocal observation. This preliminary result, that additional observations  
68 increase the first guess error, is counter-intuitive at a first glance but consistent with practical experience in  
69 weather forecasting. This finding indicates that nonlocal observations renders the LETKF unstable and it  
70 diverges dependent on properties of the observations sensitivity function. What is the role of localisation in  
71 this context? Is there a fundamental optimal relation between localisation and sensitivity function as found  
72 in [13]? The present work addresses these questions in the following sections.



**Figure 1.** Example for effect of nonlocal observations on departure statistics. Verification of local observations at three spatial positions ( $x = 1, 27$  and  $54$ ) for local observations only (open squares) and local observations and nonlocal observation (solid diamonds). The blue-coloured line sketch on the right hand side reflects the sensitivity function of the nonlocal observation with center at  $x = 40$  and the width  $r_H = 10$ ; the localisation radius is  $r_l = 10$ . Further details on the model, observations and assimilation parameters are given in section 2.7.

73 The *Methods* section introduces the essential elements of the LETKF and re-calls its analytical description  
 74 for a single observation in section 2.5. Section 2.8 provides conventional and new markers of filter  
 75 divergence that help to elucidate possible underlying divergence mechanisms. The *Results* section presents  
 76 briefly the findings, that are put into context in the *Discussion* section.

## 2 METHODS

### 77 2.1 The model

78 The storm-like Lorenz96 - model [23] is a well-established meteorological model and the present work  
 79 considers an extension by a space-dependent linear damping [24]. It is a circle network with nodes of  
 80 number  $N$ , whose node activity  $x_k(t)$  at node  $k$  and time  $t$  obeys

$$\frac{dx_k}{dt} = (x_{k+1} - x_{k-2})x_{k-1} - (1/2 + 2 \cos^4(\alpha_k \pi))x_k + I, \quad (1)$$

81 with  $k = 1, \dots, N$ ,  $x_k = x_{k+N}$  and  $\alpha_k = k/N$ . We choose  $I = 8.0$  and  $N = 80$  and the initial condition  
 82 is random with  $x_k(0) = 8.0 + \xi_k$ ,  $k \neq N/2$  and  $x_{N/2}(0) = 8.01 + \xi_{N/2}$  with the normal distributed  
 83 random variable  $\xi \sim \mathcal{N}(0, 0.01)$ . Figure 2(A) shows the model field dependent on time.

84 Typically, data assimilation techniques are applied to merge observations and solutions of imperfect  
 85 models and the true dynamics of the underlying system is not known. To illustrate the impact of nonlocal  
 86 observations, we assume (what is unrealistic in practice) that the model under consideration (1) is perfect  
 87 and hence emerging differences between observations and model equivalents do not originate in the model  
 88 error.

## 89 2.2 The Local Ensemble Transform Kalman Filter (LETKF)

90 The aim of data assimilation is to estimate a state that describes optimally both a model (or background)  
 91 state  $\mathbf{x}^b \in \mathbb{R}^N$  and corresponding observations  $\mathbf{y} \in \mathbb{R}^S$  of number  $S$ . This *analysis*  $\mathbf{x}^a \in \mathbb{R}^N$  minimizes  
 92 the cost function

$$J(\mathbf{x}) = (\mathbf{x} - \mathbf{x}^b)^t \mathbf{B}^{-1} (\mathbf{x} - \mathbf{x}^b) + \quad (2)$$

$$+ (\mathbf{y} - \hat{H}(\mathbf{x}^b))^t \mathbf{R}^{-1} (\mathbf{y} - \hat{H}(\mathbf{x}^b))$$

93 with  $\mathbf{x} \in \mathbb{R}^N$ , the background error covariance  $\mathbf{B} \in \mathbb{R}^{N \times N}$  and the observation error covariance  
 94  $\mathbf{R} \in \mathbb{R}^{S \times S}$ . The observation operator  $\hat{H} : \mathbb{R}^N \rightarrow \mathbb{R}^S$  is linear in the present work and projects a model  
 95 state into the observation space and thus links model and observations.

96 The LETKF estimates the background error covariance  $\mathbf{B}$  by background-ensemble perturbations of  
 97 number  $L$

$$\mathbf{B} \approx \frac{1}{L-1} \mathbf{X}^b (\mathbf{X}^b)^t \quad (3)$$

98 with  $\mathbf{X}^b \in \mathbb{R}^{N \times L}$ . The columns of  $\mathbf{X}^b$  are the background ensemble member perturbations  $\{\mathbf{x}^{b,l} -$   
 99  $\bar{\mathbf{x}}^b\} \forall l = 1, \dots, L$ ,  $\{\mathbf{x}^{b,l}\}$  is the set of background ensemble members and  $\bar{\mathbf{x}}^b$  is the mean over the  
 100 ensemble.

101 Then the coordinate transformation from physical space to ensemble space

$$\mathbf{x} = \bar{\mathbf{x}}^b + \mathbf{X}^b \mathbf{w} \quad (4)$$

102 describes a state  $\mathbf{x}$  in the ensemble space with new coordinates  $\mathbf{w}$  [7]. Inserting Eq. (4) into (2) yields

$$J(\mathbf{w}) = (L-1) \mathbf{w} \mathbf{w}^t + (\mathbf{y} - \bar{\mathbf{y}}^b - \mathbf{Y} \mathbf{w})^t \mathbf{R}^{-1} (\mathbf{y} - \bar{\mathbf{y}}^b - \mathbf{Y} \mathbf{w}) \quad (5)$$

103 in the new coordinate  $\mathbf{w}$ . Here  $\bar{\mathbf{y}}^b = \hat{H}(\bar{\mathbf{x}}^b) \in \mathbb{R}^S$  is the model equivalent of the background ensemble mean  
 104 in observation space and  $\mathbf{Y}^b = \hat{H}(\mathbf{X}^b)$  is the corresponding model equivalent of  $\mathbf{X}^b$ . This implies [7]

$$\hat{H}(\bar{\mathbf{x}} + \mathbf{X}^b \mathbf{w}) \approx \bar{\mathbf{y}}^b + \mathbf{Y}^b \mathbf{w},$$

105 which is valid for linear observation operators.

106 The minimization of the cost function (5) yields

$$\bar{\mathbf{w}}^a = \mathbf{A} (\mathbf{Y}^b)^t \mathbf{R}^{-1} (\mathbf{y} - \bar{\mathbf{y}}^b) \quad (6)$$

107 with

$$\mathbf{A} = \left[ (L-1) \mathbf{I} + (\mathbf{Y}^b)^t \mathbf{R}^{-1} \mathbf{Y}^b \right]^{-1}. \quad (7)$$

108 Equation (4) provides the analysis ensemble mean

$$\bar{\mathbf{x}}^a = \bar{\mathbf{x}}^b + \mathbf{X}^b \bar{\mathbf{w}}^a . \quad (8)$$

109 Then the square root filter-ansatz [7] yields the analysis ensemble members

$$\mathbf{w}^{a,l} = \bar{\mathbf{w}}^a + \mathbf{W}^{a,l} ,$$

110 where  $\mathbf{W}^{a,l}$  is the  $l$ -th column of the matrix  $\mathbf{W}^a = [(L-1)\mathbf{A}]^{1/2}$ . The square root of  $\mathbf{A}$  may be computed  
 111 by using the singular value decomposition  $\mathbf{A} = \mathbf{U}\mathbf{D}\mathbf{V}^t$  with the diagonal matrix  $\mathbf{D}$  and the eigenvector  
 112 matrices  $\mathbf{U}, \mathbf{V}$ . This yields  $\mathbf{A}^{1/2} = \mathbf{U}\mathbf{D}^{1/2}\mathbf{V}^t$ .

113 Finally the analysis ensemble members in physical space read

$$\mathbf{x}^{a,l} = \bar{\mathbf{x}}^b + \mathbf{X} \bar{\mathbf{w}}^a + \mathbf{X} \mathbf{W}^{a,l} , \quad l = 1, \dots, L , \quad (9)$$

114 see [7, 8] for more details.

115 Specifically, we have chosen  $L = 10$  ensemble members and number of observations  $S = 1$  or  $S = 2$ .

### 116 2.3 Observation data

117 In principle there are two types of observations. Local observations are measured at a single spatial  
 118 location in the system, whereas nonlocal observations are integrals over a set of spatial locations. Examples  
 119 for local observations are radiosondes measuring humidity and temperature in the atmosphere at a certain  
 120 vertical altitude and horizontal position. Typical nonlocal observations are satellite measurements capturing  
 121 the radiation in a vertical atmospheric column.

122 The present work considers observations

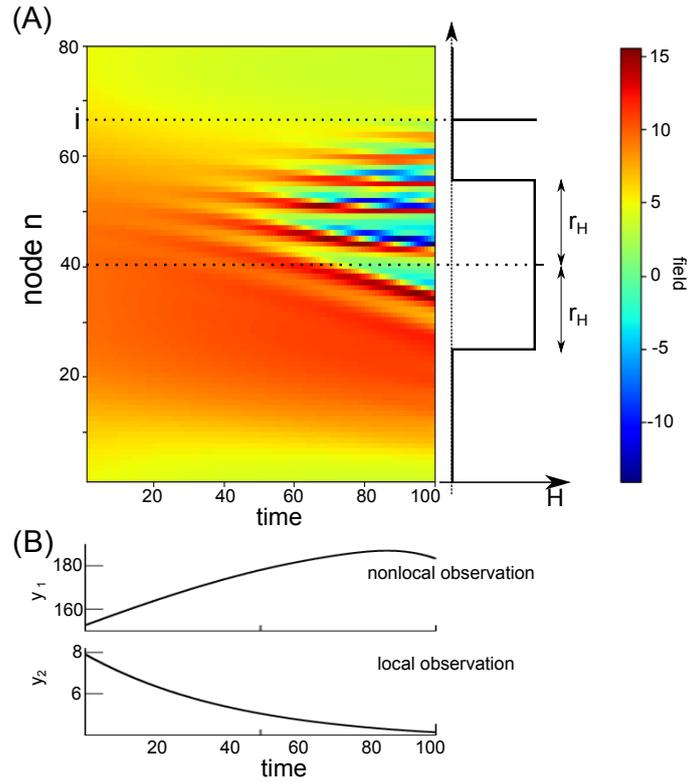
$$\mathbf{y} = \hat{H}(\mathbf{x}) + \boldsymbol{\eta} , \quad (10)$$

123 where  $\boldsymbol{\eta} \in \mathbb{S}$  is Gaussian white noise with the true variance  $\mathbf{R}_t$  and  $\hat{H}(\mathbf{x})$  is a linear observation operator  
 124  $\hat{H}(\mathbf{x}) = \mathbf{H}\mathbf{x}$ ,  $\mathbf{H} \in \mathbb{R}^{S \times N}$ . In the following, the linear operator  $\mathbf{H}$  is called sensitivity function and  
 125 we adopt this name from meteorological data assimilation of nonlocal satellite data. The present work  
 126 considers either nonlocal observations only ( $S = 1$ )

$$\begin{aligned} H_{1n} &= 1 \quad \forall n \in [N/2 - r_H; N/2[ , n \in ]N/2; N/2 + r_H] \\ H_{1n} &= 1 + 10^{-5} \quad , \quad n = N/2 \\ H_{1n} &= 0 \quad \text{otherwise} \end{aligned} \quad (11)$$

127 with sensitivity function width  $r_H$  or both observation types ( $S = 2$ )

$$\begin{aligned} H_{1n} &= 1 \quad \forall n \in [N/2 - r_H; N/2[ , n \in ]N/2; N/2 + r_H] \\ H_{1n} &= 1 + 10^{-5} \quad , \quad n = N/2 \\ H_{2n} &= 1 \quad \text{for } n = i \\ H_{km} &= 0 \quad \text{otherwise,} \end{aligned} \quad (12)$$



**Figure 2.** The model field  $\mathbf{V}$  from Eq. (1) and an illustration of the observation operator  $\mathbf{H}$  from Eq. (12) with the sensitivity function width  $r_H$ . (A) Exemplary space-time distribution of the model solution (left hand side) with the parameters  $I = 8.0$  and  $N = 80$ , the sketched position  $i$  of a local observation and a sketched sensitivity function of nonlocal observation with the center in the middle of the spatial field with radius  $r_H$ . (B) Example observations illustrating that local and nonlocal observations are scalars and may evolve differently over time. Please also note the different order of values of the two observation types.

128 where the local observation is captured at spatial location  $i$ , cf. Fig. 2 for illustration. In the subsequent  
 129 sections,  $i = N/2$  and  $r_H$  varies in the range  $1 \leq r_H \leq 10$ . Please note that  $r_H = 1$  approximates a  
 130 local observation. Moreover, in the following a grid point whose activity contributes to an observation is  
 131 called an *observed grid point* and all others are called *unobserved grid points*. Mathematically, observed  
 132 (unobserved) grid points exhibit  $H_{nk} \neq 0$  ( $H_{nk} = 0$ ).

In this work, a single partial study considers a smooth sensitivity function instead of the boxcar function described above. Then the sensitivity function is the Gaspari-Cohn function  $GC(n, r_H/2)$  [29] in the interval  $-r_H \leq n \leq r_H$ , which approximates the Gaussian function by a smooth function with finite support  $2r_H$

$$H_{1n} = GC(n, r_H) \quad \forall n \in [-r_H; r_H]$$

$$H_{1n} = 0 \quad \text{otherwise .}$$

133 The observations  $\mathbf{y}(t_n)$ ,  $1, \dots, T$  at  $T$  time instances (cf. Eq.(10)) obey the model (1) and Eq. (10)  
 134 with the observation operator (11) or (12). In a large part of the work, we have assumed zero observation  
 135 error  $\mathbf{R}_t = 0$ , i.e. observations are perfect in the sense that they reflect the underlying perfect model, cf.  
 136 section 2.1. We take the point of view that we do not know that the model and observations are perfect and  
 137 hence we guess  $\mathbf{R}$  as it is done in cases where models and observations are not perfect.

138 This approach has been taken in most cases in the work. Since, however, this implicit filter error may already  
 139 contribute to a filter instability or even may induce it, a short partial study has assumed perfect knowledge of  
 140 the observation error. To this end, in this short partial study we have assumed  $(\mathbf{R}_t)_{jj} = 0.1$ ,  $j = 1, \dots, S$   
 141 and perfect knowledge of this error, i.e.  $\mathbf{R} = \mathbf{R}_t$ .

142 Although techniques have been developed to estimate  $\mathbf{R}$  adaptively [25], we do not employ such a scheme  
 143 for simplicity.

144

## 145 2.4 Localisation

146 In the LETKF, the background covariance matrix  $\mathbf{B}$  is expressed by  $L$  ensemble members, cf. Eq. (3),  
 147 and it is rank-deficit for  $L \ll N$ . This leads to spurious correlations in  $\mathbf{B}$ . Spatial localisation in ensemble  
 148 Kalman filters has been found to be beneficial [16, 26, 27, 28] in this context. The LETKF as defined by  
 149 Hunt et al. [7] performs the localization in observation space. In detail, Hunt et al. [7] proposed to localize  
 150 by increasing the observation error in matrix  $\mathbf{R}$  dependent on the distance between the analysis grid point  
 151 and observations. The present implementation follows this approach.

152 The observation error matrix  $\mathbf{R}$  is diagonal, i.e. observation errors between single observations are  
 153 uncorrelated. Then at each grid point  $i$  the localisation scheme considers observations  $y_n$  at location  $j$   
 154 only if the distance between location  $i$  and  $j$  does not exceed the localisation radius  $r_l$ . Then the error of  
 155 observation  $n$  is  $R_{nn} = R_{nn}^0 / \rho_{ij}$ , where  $\rho_{ij} = GC(d_{ij}, r) + \varepsilon$  for  $d_{ij} \leq r_l$  is the weighting function with  
 156 the Gaspari-Cohn function  $GC(d, r)$  [29],  $\varepsilon > 0$  is a small constant ensuring a finite observation error and  
 157  $d_{ij}$  is the spatial distance between  $i$  and  $j$ . The Gaspari-Cohn function approximates a Gaussian function  
 158 with standard deviation  $r\sqrt{3/10}$  by a polynomial with finite support. The parameter  $2r = r_l$  is the radius  
 159 of the localisation function with  $0 \leq GC(z, r) \leq 1$ ,  $0 \leq z \leq r_l$ . Consequently the observation error takes  
 160 its minimum  $R_{nn}^0$  at distance  $d_{ij} = 0$  and increases monotonously with distance to its maximum  $R_{nn}^0 / \varepsilon$  at  
 161  $d_{ij} = r_l$ . In the present implementation, we use  $\varepsilon = 10^{-7}$  and observation errors  $R_{11}^0 = 0.1$  for a single  
 162 nonlocal observation  $S = 1$  and  $R_{nn}^0 = 0.1\delta_{nm}$ ,  $n = 1, 2$  for local and nonlocal observation with  $S = 2$ .

163 The observation error close to the border of the localisation area about a grid point  $i$  is large by definition  
 164  $R_{nn} = R_{nn}^0 / (GC(d \rightarrow r_l/2, r_l/2) + \varepsilon)$ . In numerical practice, the assimilation effect of large values  
 165  $R_{nn} > R_{nn}^0 / GC_{\text{low}}$  is equivalent for some distances from the grid point  $i$  in a reasonable approximation if  
 166  $GC_{\text{low}}$  is low enough. By virtue of the monotonic decrease of  $GC(d, r_l/2)$  with respect to distance  $d \geq 0$ ,  
 167 this yields the condition  $GC(r_l \geq d \geq r_c, r_l/2) < GC_{\text{low}}$ . In other words, for distances  $d$  larger than a  
 168 corrected localisation radius  $r_c$ , the observation errors  $R_{nn}$  are that large that observations at such distances  
 169 do poorly contribute to the analysis. For instance, if  $GC_{\text{low}} = 0.01$ , then  $r_l = 5 \rightarrow r_c = 3$ ,  $r_l = 10 \rightarrow$   
 170  $r_c = 7$  and  $r_l = 15 \rightarrow r_c = 11$ . It is important to note that this corrected localisation radius depends on the  
 171 width of the Gaspari-Cohn function and thus on the original localisation radius  $r_l$ , i.e.  $r_c = r_c(r_l)$ . In most  
 172 following study cases results are given for original localisation radii  $r_l$ , while the usage of the corrected  
 173 localisation radius is stated explicitly. [The existence of a corrected localisation radius  \$r\_c\$  illustrates the insight, that there is not a single optimal localisation radius for smooth localisation functions but a certain range of equivalent localisation radii. For non-smooth localisation functions with sharp edges, e.g. a boxcar function, this variability would not exist.](#)

177 The present work considers primarily nonlocal observations. Since these are not located at a single spatial  
 178 site, it is non-trivial to include them in the LETKF that assumes a single observation location. To this end,  
 179 several previous studies have suggested corresponding approaches [30, 31, 24, 32, 33, 34, 35, 36, 13]. A  
 180 reasonable approximation for the spatial location of a nonlocal observation is the location of the maximum

181 sensitivity [37, 10], i.e.  $\max_n H_{kn}$  of nonlocal observation  $k$ . Although this approximation has been shown  
 182 to yield good results, it introduces a considerable error for broad sensitivity functions, i.e.  $r_H$  is large. In  
 183 fact, this localisation scheme introduces an additional contribution to the observation error. The present  
 184 implementation considers this definition. This results in the localisation of the nonlocal observation at grid  
 185 point  $i = N/2$ .

## 186 2.5 LETKF for a single observation

187 In a large part of this work, we consider a single observation with  $S = 1$ . The subsequent paragraphs  
 188 show an analytical derivation of the ensemble analysis mean and the analysis members, whose terms are  
 189 interpreted in the Results section.

190 Considering the localisation scheme described above, at the model grid point  $i$  the analysis ensemble  
 191 mean (8) reads

$$\bar{x}_i^a = \bar{x}_i^b + (\mathbf{X}_i \mathbf{A}_i \mathbf{Y}^t) (y_0 - \bar{y}^b) / R_i \quad (13)$$

192 where  $\mathbf{Y} \in \mathbb{R}^L$  is a row vector with  $Y_k = Y_{1k}^b$ , with the row vector  $\mathbf{X}_i \in \mathbb{R}^L$ ,  $(\mathbf{X}_i)_k = X_{ik}$  and

$$\mathbf{A}_i = [ (L-1)\mathbf{I} + \mathbf{Y}^t \mathbf{Y} / R_i ]^{-1} . \quad (14)$$

193 The term  $R_i = R_{11}^0 / \rho_{i(N/2)}$  denotes the weighted observation error used at grid point  $i$ , when the  
 194 observation is located at  $j = N/2$ , and  $R_{11}^0$  is the error of observation  $y_1$ .

195 Now utilising the Woodbury matrix identity [38]

$$\begin{aligned} & (\mathbf{B} + \mathbf{UCV})^{-1} = \\ & \mathbf{B}^{-1} - \mathbf{B}^{-1} \mathbf{U} (\mathbf{C}^{-1} + \mathbf{V} \mathbf{B}^{-1} \mathbf{U})^{-1} \mathbf{V} \mathbf{B}^{-1} \end{aligned}$$

196 for real matrices  $\mathbf{B} \in \mathbb{R}^{n \times n}$ ,  $\mathbf{U} \in \mathbb{R}^{n \times k}$ ,  $\mathbf{C} \in \mathbb{R}^{k \times k}$  and  $\mathbf{V} \in \mathbb{R}^{k \times n}$  with  $n, k \in \mathbb{N}$ , Eq. (14) reads

$$\begin{aligned} \mathbf{A}_i &= \frac{1}{L-1} \mathbf{Q}_i \\ \mathbf{Q}_i &= \mathbf{I} - \frac{1}{(L-1)R_i + y^2} \mathbf{Y}^t \mathbf{Y} , \end{aligned} \quad (15)$$

197 where  $y = \sqrt{\mathbf{Y} \mathbf{Y}^t} \in \mathbb{R}$  is a scalar. Inserting (15) into (13), the analysis ensemble mean is

$$\bar{x}_i^a = \bar{x}_i^b + \mathbf{X}_i \mathbf{Y}^t \alpha_i \quad (16)$$

198 with

$$\alpha_i = \frac{y_0 - \bar{y}^b}{(L-1)R_i + y^2} .$$

199 Since  $R_{N/2} = R_{11}^0$  and  $R_{N/2 \pm r_i} = R_{nn}^0 / \varepsilon = 10^7 R_{nn}^0$ ,  $\alpha_i$  takes its maximum at the observation location  
 200 and is very small when the observation is at the localisation border. This means that  $\bar{x}_i^a \approx \bar{x}_i^b$  at the border  
 201 of the localisation area.

202 Now let us focus on the ensemble members. Equations (16) and (9) give the analysis ensemble members  
203 at grid point  $i$

$$x_i^{a,l} = \bar{x}_i^b + \mathbf{X}_i \mathbf{Y}^t \alpha_i + \sqrt{L-1} \mathbf{X}_i \left( \sqrt{\mathbf{A}_i} \right)_l, \quad (17)$$

204 where  $(\sqrt{\mathbf{A}_i})_l$  is the  $l$ -th column of matrix  $\sqrt{\mathbf{A}_i} = \sqrt{\mathbf{Q}_i} / \sqrt{L-1}$ .  
205 The singular value decomposition serves as a tool to compute

$$\sqrt{\mathbf{Q}_i} = \mathbf{U} \sqrt{\mathbf{D}} \mathbf{U}^t, \quad (18)$$

206 where  $\sqrt{\mathbf{D}} \in \mathbb{R}^{L \times L}$  is diagonal and its matrix elements are the eigenvalues of  $\mathbf{Q}$ . The columns of matrix  
207  $\mathbf{U} \in \mathbb{R}^{L \times L}$  are the normalised eigenvectors of  $\mathbf{Q}$ . Then Eq. (7) yields

$$\begin{aligned} \mathbf{Q}_i \mathbf{Y}^t &= \left( \mathbf{I} - \frac{1}{(L-1)R_i + y^2} \mathbf{Y}^t \mathbf{Y} \right) \mathbf{Y}^t \\ &= \mathbf{Y}^t - \frac{y^2}{(L-1)R_i + y^2} \mathbf{Y}^t \\ &= \frac{(L-1)R_i}{(L-1)R_i + y^2} \mathbf{Y}^t \\ &= \lambda_i \mathbf{Y}^t, \end{aligned}$$

208 i.e.  $\mathbf{Y}^t$  is an eigenvector of  $\mathbf{Q}_i$  with eigenvalue  $0 < \lambda_i < 1$ . By virtue of the properties of  $R_i$ ,  $\lambda_i$  takes its  
209 minimum at the observation location at  $i = N/2$  and it is maximum at the localisation border.  
210 The remaining eigenvectors of number  $L-1$  are  $\mathbf{v}_n \perp \mathbf{Y}$ ,  $n = 1, \dots, L-1$  with unity eigenvalue since

$$\begin{aligned} \mathbf{Q}_i \mathbf{v}_n &= \left( \mathbf{I} - \frac{1}{(L-1)R_i + y^2} \mathbf{Y}^t \mathbf{Y} \right) \mathbf{v}_n \\ &= \mathbf{v}_n - \frac{1}{(L-1)R_i + y^2} \mathbf{Y}^t \underbrace{\mathbf{Y} \mathbf{v}_n}_{=0} \\ &= \mathbf{v}_n. \end{aligned}$$

211 Hence  $\mathbf{U} = (\mathbf{Y}^t / \|\mathbf{Y}\|, \mathbf{v}_1, \dots, \mathbf{v}_{L-1})$  and  $\sqrt{\mathbf{D}} = \text{diag}(\sqrt{\lambda_i}, 1, \dots)$  and, after inserting into Eq. (18) and  
212 lengthy calculations

$$(\sqrt{\mathbf{Q}_i})_{kl} = \sqrt{\lambda_i} \frac{Y_k Y_l}{y^2} + \sum_{n=1}^{L-1} (\mathbf{v}_n)_k (\mathbf{v}_n)_l. \quad (19)$$

213 This leads to

$$x_i^{a,l} = \bar{x}_i^b + \mathbf{X}_i \mathbf{Y}^t \alpha_i + \mathbf{X}_i \mathbf{Y}^t \frac{\sqrt{\lambda_i}}{y^2} Y_l + \sum_{n=1}^{L-1} \mathbf{X}_i \mathbf{v}_n (\mathbf{v}_n)_l. \quad (20)$$

## 214 2.6 Additive covariance inflation

215 The ensemble Kalman filter underestimates the forecast error covariance matrix due to the limited  
 216 ensemble size [39]. This problem is often addressed by covariance inflation [40, 41, 26]. The present work  
 217 implements additive covariance inflation [42]. The ensemble perturbations  $\mathbf{X}^b$  in (3) are modified by white  
 218 Gaussian additive noise  $\Gamma \in \mathbb{R}^{N \times L}$

$$\mathbf{X}_{\text{add}}^b = \mathbf{X}^b + \Gamma .$$

219 with matrix elements  $\Gamma_{ij} \sim \mathcal{N}(0, f_{\text{add}}^2)$  and the inflation factor  $f_{\text{add}} = 0.1$ .

## 220 2.7 Numerical experiments

221 The present study investigates solutions  $\mathbf{x}(t)$  of model (1) and [Eq. \(10\) provides the](#) observations  $\mathbf{y}(t)$ .  
 222 This is called the nature run. In the filter cycle, the initial analysis values are identical to the initial values  
 223 of the nature run and the underlying filter model is the true model (1). In the forecast step, the model is  
 224 advanced with time step  $\Delta t = 10^{-3}/12$  for 100 time steps applying a 4th-order Runge-Kutta integration  
 225 scheme. According to [23], the duration of one forecast step corresponds to 1 hour which is also the  
 226 time between two successive observations. The analysis update is instantaneous. In an initial phase, the  
 227 model evolves freely for 50 forecast steps to avoid possible initial transients. Then, the LETKF estimates  
 228 the analysis ensemble according to section 2.2 during 200 cycles if not stated otherwise. One of such  
 229 a numerical simulation is called a trial in the following. Each trial assumes identical initial ensemble  
 230 members and the only difference in trials results from the additive noise in additive covariance inflation, cf.  
 231 section 2.6.

232 [By virtue of the primarily numerical nature of the present work, it is mandatory to vary certain parameters,](#)  
 233 [such as perturbations to the observations or the factor of additive inflation. For instance, the data assimilation](#)  
 234 [results in Figure 1 are based on model \(1\), 3 local and 1 nonlocal observation. This corresponds to the](#)  
 235 [observation operator  \$\hat{H}\$  with the sensitivity function](#)

$$\begin{aligned} H_{1n} &= \delta_{nn_1} \\ H_{2n} &= \delta_{nn_2} \\ H_{3n} &= \delta_{nn_3} \\ H_{4n} &= 1 \quad \forall n \in [N/2 - r_H; N/2 + r_H] \end{aligned}$$

236 with  $r_H = 10$  and  $n_1 = 1$ ,  $n_2 = 27$ ,  $n_3 = 54$ . The localisation radius is identical to the sensitivity function  
 237  $r_l = r_H$  and data assimilation is performed during 250 filter cycles with an initial phase of 50 forecast steps.  
 238 For stabilisation reasons, we have increased the model integration time step to  $\Delta t = 10^{-2}/12$  but reduced  
 239 the number of model integrations to 10 steps, cf. [19], thus essentially retaining the time interval between  
 240 observations. Other parameters are identical to the standard setting described in the previous sections.

241 [As mentioned above, typically the measurement process is not known in all details. For instance, the](#)  
 242 [observation error is assumed to be  \$R = 0.1\$  for the nonlocal observations, whereas the true model exhibits](#)  
 243 [noise-free observations with  \$R\_t = 0\$ . This is the valid setting for all simulations but few set of trials shown](#)  
 244 [in Fig. 5. In a set of experiments \(Fig. 5\(solid, dashed and dashed-dotted line\)\), observations are noisy with](#)  
 245 [noise perturbation variance 0.1 and hence  \$R\_t = R = 0.1\$ . Moreover, the additive inflation factor is chosen](#)  
 246 [to  \$f\_{\text{add}} = 0.1\$  but in two single sets of experiments \(cf. Fig. 5\(dashed and dashed-dotted line\)\), where](#)  
 247 [the additive inflation factor is  \$f\_{\text{add}} = 0.05\$ . In addition, the weighting function of nonlocal observations is a boxcar window function with](#)

248 sharp borders but in a single set of experiments, where the weighting function is a smooth Gaspari-Cohn  
 249 function, cf. Fig. 5(dashed-dotted line).

250 The verification measures bias and RMSE are computed for the local observations only according to  
 251 Eqs. (21),(22).

## 252 2.8 Divergence criteria and verification

253 The Kalman filter may diverge for several reasons [26, 6, 43], such as model error, insufficient sampling  
 254 of error covariance or high condition number of observation operators [17, 44]. Especially the latter has  
 255 been shown to be able to trigger catastrophic filter divergence of the ensemble Kalman filter exhibiting a  
 256 diverging forecasts in model state space [19, 21, 20]. This divergence type exhibits a magnitude increase of  
 257 model variables to machine infinity in finite time. The present implementation detects catastrophic filter  
 258 divergence and stop the numerical simulation when the maximum absolute value of any single ensemble  
 259 member exceeds a certain threshold  $|x_k^{b,l}| > 10^{10}$ ,  $k \in [1; N]$ ,  $l \in [1; L]$ .

260 The present work focuses primarily on a non-catastrophic filter divergence type showing a strong increase  
 261 of the innovation magnitude to values much larger than the observation equivalent of the attractor. This  
 262 divergence may be temporally intermittent with finite duration. Since this intermittent innovation divergence  
 263 results in increased first guess departures and hence worsens forecasts, it is important to detect these  
 264 divergences and control them. By definition the innovation process diverges if  $\max_{l,k} |[\mathbf{y}_n - \mathbf{H}\mathbf{x}^{b,l}]_k| > \sigma_{th}$   
 265 for any observation  $n$  with  $\sigma_{th} = 1000\sqrt{R_{nn}^0}$ . Then the numerical simulation is stopped. The time of filter  
 266 divergence is called  $T_b$  in the following. This criterion for innovation divergence is hard: if the innovation  
 267 reaches the threshold  $\sigma_{th}$ , then innovation divergence occurs. The corresponding divergence rate  $\gamma$  is the  
 268 ratio between the number of **divergent and non-divergent** trials. For instance, for  $\gamma = 1$  all numerical trials  
 269 diverge whereas  $\gamma = 0$  reflect stability in all numerical trials.

270 Moreover, it is possible that  $|[\mathbf{y}_n - \mathbf{H}\mathbf{x}^{b,l}]_k|$  grows intermittently but does not reach the divergence  
 271 threshold. The first guess departure bias

$$\text{bias} = \frac{1}{TS} \sum_{k=1}^T \sum_{n=1}^S [\mathbf{y}(t_k)]_n - [\mathbf{H}\bar{\mathbf{x}}^b(t_k)]_n \quad (21)$$

272 and the corresponding root mean-square error

$$\text{RMSE} = \frac{1}{TS} \sum_{k=1}^T \sum_{n=1}^S (\mathbf{y}_n(t_k) - [\mathbf{H}\bar{\mathbf{x}}^b(t_k)]_n)^2 \quad (22)$$

273 quantify the forecast error in such trials. For a single observation,  $\mathbf{y} \rightarrow y_o$ . Larger values of bias RMSE  
 274 indicate larger innovation values.

275 To quantify filter divergence, Tong et al. [18] have proposed the statistical measure

$$\Theta_n = \sqrt{\frac{1}{L} \sum_{l=1}^L (\mathbf{y}(t_n) - \mathbf{H}\mathbf{x}^{b,l}(t_n))^t (\mathbf{y}(t_n) - \mathbf{H}\mathbf{x}^{b,l}(t_n))}$$

276 and

$$\Xi_n = \left\| \frac{1}{L-1} \sum_{l=1}^L \mathbf{X}_o(t_n) \otimes \mathbf{X}_u(t_n) \right\|$$

277 at time  $t_n$ , where the norm is defined by  $\|\mathbf{Z}\| = \sum_{n,m} |Z_{nm}|^2$  for any matrix  $\mathbf{Z}$  and  $Z_{nm}$  are the  
 278 corresponding matrix elements. The quantity  $\Theta_n$  represents the ensemble spread in observation space and  
 279  $\Xi_n$  is the covariation of observed and unobserved ensemble perturbations assuming local observations.  
 280 Large values of  $\Xi$  indicates catastrophic filter divergence as pointed out in [18, 20]. This definition may also  
 281 apply to nonlocal observations, cf. section 2.5, although its original motivation assumes local observations.  
 282 An interesting feature to estimate the degree of divergence is the time of maximum ensemble spread  $T_\Theta$   
 283 and the time of maximum covariation of observed and unobserved ensemble perturbations  $T_\Xi$ :

$$\begin{aligned} T_\Theta &= \arg \max_n \Theta_n \\ T_\Xi &= \arg \max_n \Xi_n . \end{aligned} \quad (23)$$

284

285 Moreover, previous studies have pointed out that catastrophic filter divergence in ensemble Kalman  
 286 filter implies alignment of ensemble members. This may also represent an important mechanism in  
 287 non-catastrophic filter divergence. The new quantity

$$p_{a,u} = \frac{n_{a,u}}{L(L-1)/2} \quad (24)$$

288 is the probability of alignment and unalignment, where  $n_a$  is the number of aligned ensemble member  
 289 perturbation pairs  $(\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b)$ ,  $(\mathbf{x}^{b,k} - \bar{\mathbf{x}}^b)$  for which

$$\cos \beta_{lk} = \frac{(\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b)^t (\mathbf{x}^{b,k} - \bar{\mathbf{x}}^b)}{\|\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b\| \|\mathbf{x}^{b,k} - \bar{\mathbf{x}}^b\|} \geq 0.5$$

290 and  $n_u$  is the number of ant-aligned member pairs with

$$\cos \beta_{lk} = \frac{(\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b)^t (\mathbf{x}^{b,k} - \bar{\mathbf{x}}^b)}{\|\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b\| \|\mathbf{x}^{b,k} - \bar{\mathbf{x}}^b\|} \leq -0.5$$

291  $\forall l \neq k, l, k = 1, \dots, L$ . The alignment (anti-alignment) condition  $\cos \beta_{lk} > 0.5$  ( $\cos \beta_{lk} < -0.5$ ) implies  
 292  $-60^\circ \leq \beta_{lk} \leq 60^\circ$  ( $120^\circ \leq \beta_{lk} \leq 240^\circ$ ). Please note that  $0 \leq p_{a,u} \leq 1$  and the larger  $p_a$  ( $p_u$ ) the more  
 293 ensemble members are aligned (anti-aligned) to each other.

294 Considering the importance of member alignment to each other for catastrophic divergence, it may  
 295 be interesting to estimate the alignment degree of background member perturbation with the analysis  
 296 increments  $\mathbf{x}^{a,l} - \mathbf{x}^{b,l}$  by

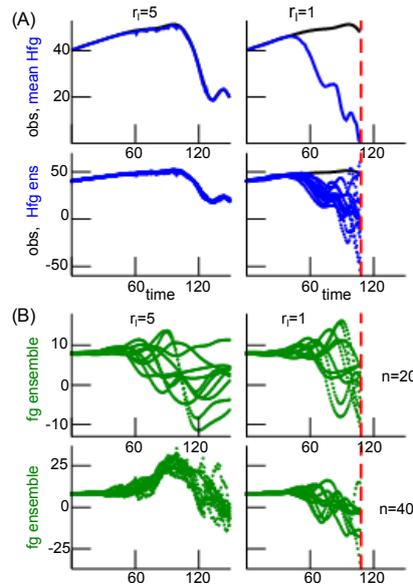
$$\cos \alpha_l = \frac{(\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b)^t (\mathbf{x}^{a,l} - \mathbf{x}^{b,l})}{\|\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b\| \|\mathbf{x}^{a,l} - \mathbf{x}^{b,l}\|}, \quad l = 1, \dots, L. \quad (25)$$

297 The term  $\mathbf{x}^{a,l} - \mathbf{x}^{b,l}$  is the analysis ensemble member perturbation from the background members and  
 298  $\mathbf{x}^{b,l} - \bar{\mathbf{x}}^b$  is the direction of the background member perturbation. If  $\cos \alpha_l \rightarrow 1$  ( $\cos \alpha_l \rightarrow -1$ ) the analysis

299 ensemble members point into the same (opposite) direction as the background ensemble members. In  
300 addition,

$$q_a = \frac{n_a}{L} \quad , \quad q_u = \frac{n_u}{L} \quad (26)$$

301 are the percentages of aligned and anti-aligned ensemble members for which  $\cos \alpha_l > 0.5$  (of number  $n_a$ )  
302 and  $\cos \alpha_l < -0.5$  (of number  $n_u$ ), respectively.



**Figure 3.** Temporal solutions of the filter process with  $r_H = 5$  with two different localisation radii  $r_l$ . (A) Comparison of observations (black line) and model equivalents of the ensemble mean  $\mathbf{y}^b$  (top row, solid blue line) and the ensemble members  $\mathbf{y}^{(b,l)}$  (bottom row, dotted blue line). The time represents the number of analysis steps. (B) Ensemble members in model space at the single spatial location  $n = 20$  (shown in top panel), i.e. outside the observation area with  $H_{1n} = 0$ , and at the single spatial location  $n = 40$ , i.e. in the center of the observation area (shown in bottom panel).

### 3 RESULTS

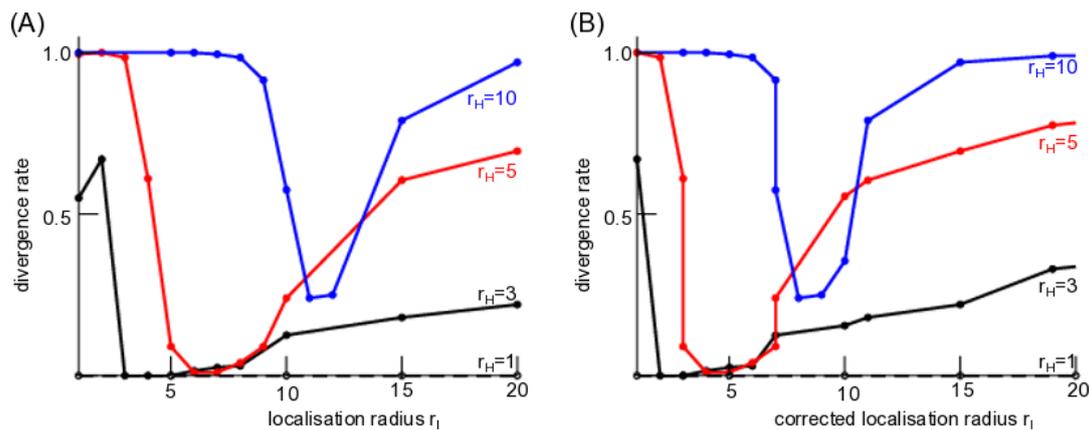
303 The stability of the ensemble Kalman filter depends heavily on the model and the nature of observations.  
 304 To gain some insight into the effect of nonlocal observations, the present work considers primarily nonlocal  
 305 observations only (section 3.1). Then the last section (3.2) shows briefly the divergence rates in the  
 306 presence of both local and nonlocal observations.

#### 307 3.1 Nonlocal observations

308 The subsequent sections consider nonlocal observations only and show how they affect the filter stability.  
 309 To this end, the first studies are purely numerical and are complemented by an additional analytical study.

#### 310 Numerical results

311 In order to find out how the choice of localisation radius  $r_l$  affects the stability of the LETKF, a large  
 312 number of numerical experiments assist to investigate statistically under which condition the filter diverges.  
 313 Figure 3 shows the temporal evolution of the background  $\mathbf{x}^b$  and the model equivalents in observation  
 314 space  $\mathbf{y}^b$  for two different localisation radii. In Fig. 3(A) observations (black line) are very close to model  
 315 equivalents (blue lines) for identical localisation and sensitivity function width, i.e.  $r_l = r_H$ . Conversely,  
 316 observations and model-equivalents diverge after some time for  $r_l \neq r_H$ . This is visible in the ensemble  
 317 mean (Fig. 3(A), top row) and the single ensembles (Fig. 3(A), bottom row). The different filter behavior  
 318 can be observed in model space as well, but there it is less obvious, cf. Fig. 3(B). The ensemble members at  
 319 spatial location  $n = 40$  are located in the center of the observation area. They exhibit a rather small spread  
 320 around the ensemble mean for  $r_l = r_H$ , whereas the ensemble spread is larger for  $r_l \neq r_H$ . The ensemble  
 321 at  $n = 20$  is outside the observation area and thus is not assimilated. There, the ensemble in  $r_l = r_H$  and  
 322  $r_l \neq r_H$  are close to each other.



**Figure 4.** Stability of the LETKF of nonlocal observations dependent on the sensitivity function width  $r_H$  and the localisation radius  $r_l$ . The divergence rate  $\gamma$  is defined in section 2.8. (A) with original localisation radius  $r_l$ . (B) with corrected localisation radius  $r_c$  and  $GC_{\text{low}} = 0.01$ . Here, the observations are noise-free with  $R_t = 0$  but the chosen observation error is assumed to  $R = 0.1 \neq R_t$  due to lack of knowledge of this true value.

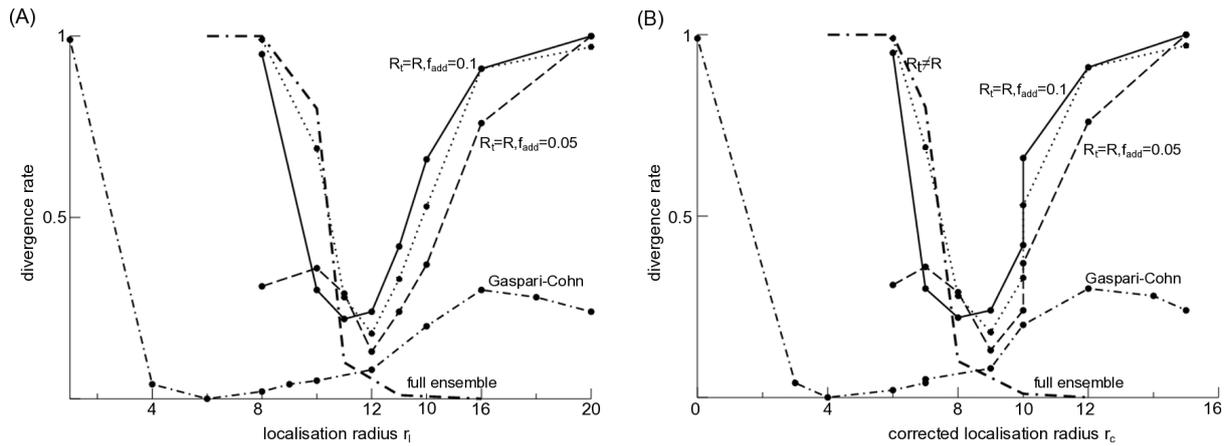
323 This result can be generalised to a larger number of localisation and sensitivity function widths, cf. Fig. 4.  
 324 For the smallest sensitivity function width and thus the smallest observation area with  $r_H = 1$ , no filter  
 325 process diverges for a large range of localisation radii  $r_l$ , i.e. the LETKF is stable (dashed black line in  
 326 Fig. 4). This case  $r_H = 1$  corresponds to local observations. Now increasing the observation area with  
 327  $r_H > 1$ , the filter may diverge and its divergence rate  $\gamma$  depends on the localisation radius. We observe that  
 328 the filter diverges least when the localisation radius is close to the sensitivity function width. These findings  
 329 hold true for both the original localisation radius and the corrected radius  $r_c$ , cf. Methods section 2.4 and  
 330 Fig. 4(A) and (B). Moreover, the filter does not exhibit catastrophic divergence before the background  
 331 reaches its divergence threshold.

332 These results hold also true if observations are subjected to additive noise and the observation error  
 333 is chosen to the true value, cf. Fig. 5(solid line) and if additive inflation is chosen to a lower value  
 334 (Fig. 5(dashed line)). Similarly to Fig. 4, the divergence rate is minimum if the sensitivity function width is  
 335 close to the original (Fig. 5(A)) or corrected (Fig. 5(B)) localisation radius  $r_l$ . The situation is different if  
 336 the sensitivity function is not a non-smooth boxcar function as in the majority of the studies but a smooth  
 337 Gaspari-Cohn function. Then the divergence rate is still minimum but the corresponding localisation radius  
 338 of this minimum is much smaller than  $r_h$ , cf. dotted-dashed line in Fig. 5.

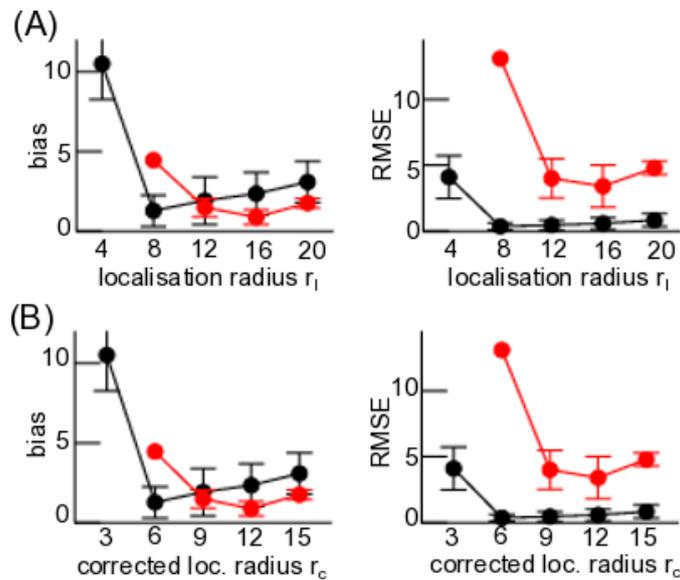
339 All these results consider the realistic case of a small number of ensemble members  $L \ll N$ . Nevertheless,  
 340 it is interesting to raise the question how these results depend on the ensemble size. Figure 5(bold dotted-  
 341 dashed line) indicates that a full ensemble with  $L = 80$  removes the minimum with maximum divergence  
 342 rate for  $r_l < r_H$  and full stability for  $r_l > r_H$ .

343 The divergence criterion is conservative with a hard threshold and trials with large but sub-threshold  
 344 innovations, i.e. with innovations that do not exceed the threshold, are not detected as being divergent.  
 345 Nevertheless to quantify intermittent large innovations in the filter, Figure 6 shows the bias and RMSE  
 346 of trials whose innovation process do not reach the divergence threshold. We observe minimum bias and  
 347 RMSE for original localisation radii  $r_l$  that are similar to the sensitivity function width  $r_H$  (Fig. 6(A)). For  
 348 corrected localisation radii  $r_c$  and  $r_H$  agree well at minimum bias and RMSE, cf. Fig. 6(B).

349 Now understanding that localisation radii  $r_l \neq r_H$  may destabilize the filter, the question arises where this  
 350 comes from and which mechanisms may be responsible for the innovation divergence. Figure 7 illustrates

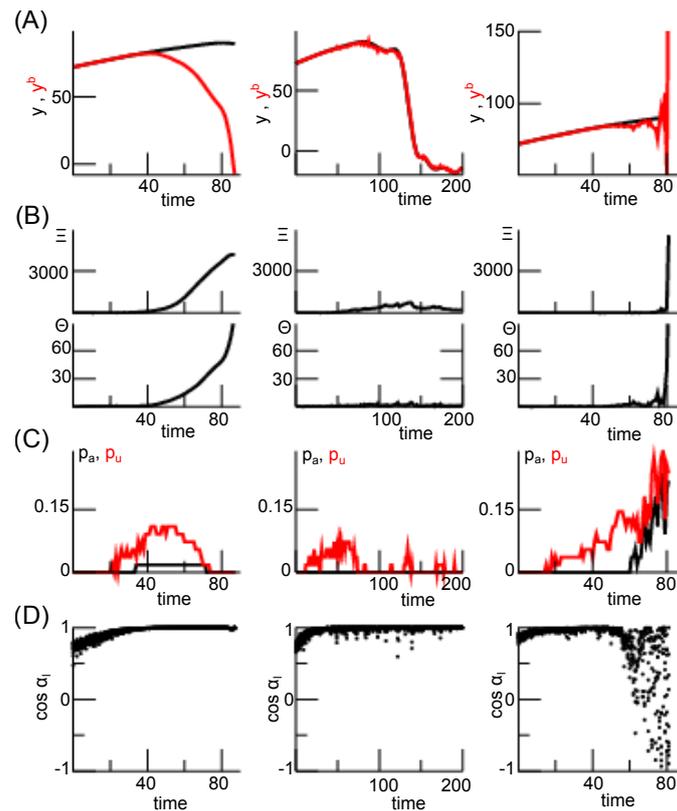


**Figure 5.** LETKF stability for different parameters and  $r_H = 10$ . The solid line denotes the divergence rate  $\gamma$  if the true observation error  $R_t = 0.1$  is known, i.e.  $R = R_t$ , and the inflation rate is  $f_{add} = 0.1$ ; the dashed line denotes the divergence rate for lower inflation rate  $f_{add} = 0.05$ , otherwise identical to the solid line case; the dotted-dashed line marks results identical to the dashed line case but with a smooth Gaspari-Cohn sensitivity function. The dotted line is taken from Fig. 4 for comparison ( $R_t = 0, R = 0.1$ ) and the bold dotted-dashed line represents the results with a full ensemble  $L = 80$ , otherwise identical to the dotted line-case. (A) original localisation radius  $r_l$ . (B) corrected localisation radius  $r_c$  with  $GC_{low} = 0.01$ .



**Figure 6.** First guess departure statistics of trials that do not reach the divergence threshold. Here  $r_H = 5$  (black) and  $r_H = 10$  (red). (A) original localisation radius  $r_l$ . (B) corrected localisation radius  $r_c$  with  $GC_{low} = 0.01$ . All statistical measures are based on 100 trials.

351 various statistical quantities for three exemplary trials. These quantities have been proposed to reflect  
 352 or explain divergence. The innovation-based measure  $\Theta_n$  diverges (Fig. 7(B)) when the filter diverges  
 353 (Fig. 7(A)) for  $r_l < r_H$  and  $r_l \gg r_h$ , whereas  $\Theta_n$  remains finite for  $r_l \approx r_H$ . Interestingly, for  $r_l < r_H$  a  
 354 certain number of ensemble members align and anti-align intermittently but do not align in the instance of  
 355 divergence (Fig. 7(C)). In the case of similar localisation radius and sensitivity function width, a similar  
 356 number of ensemble members align and anti-align but the filter does not diverge. Conversely, for  $r_l \gg r_H$   
 357 ensemble members both align and anti-align while the filter diverges. These results already indicate a  
 358 different divergence mechanism for  $r_l \leq r_H$  and  $r_l > r_H$ . Accordingly, for  $r_l < r_H$  and  $r_l \approx r_H$  background



**Figure 7.** Various measures reflecting stability of the LETKF dependent on the localisation radius  $r_l$  in single trials. (A) observation  $y_o$  (black) and model equivalent  $\mathbf{H}\bar{\mathbf{x}}^b$  (red). (B) Statistical quantities  $\Xi_n$  (top) and  $\Theta_n$  (bottom), for definition see section 2.8. (C) The probability of ensemble member alignment according to Eq. (24) for aligned (black) and anti-aligned (red) members. (D) Statistical estimate of alignment between ensemble members and  $\mathbf{x}^a - \mathbf{x}^b$  according to Eq. (25). The different localisation radii are  $r_l = 1$  (left panel),  $r_l = 6$  (center panel) and  $r_l = 20$  (right panel) with the sensitivity function width  $r_H = 5$ .

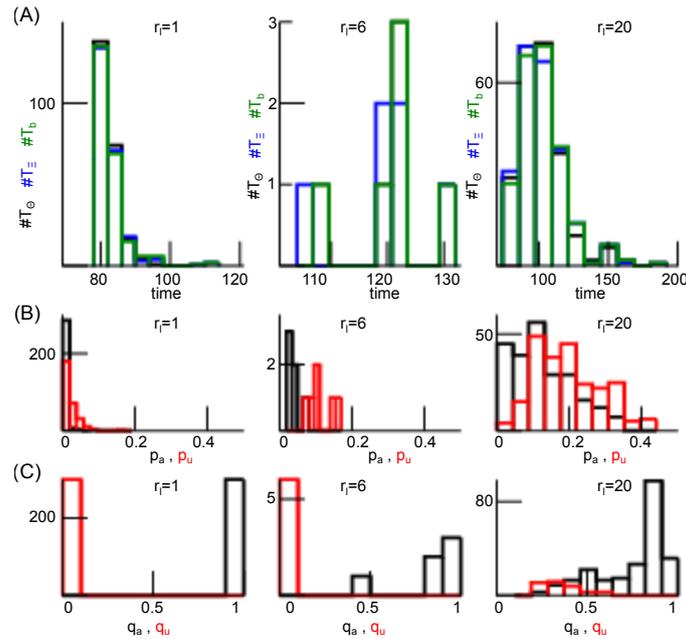
359 member perturbations align with the analysis member perturbations with  $\cos \alpha_l \rightarrow 1$  (Fig. 7(D)), whereas  
 360  $\cos \alpha_l$  fluctuates between 1 and  $-1$  for  $r_l \gg r_H$  while diverging.

361 Figure 8(A) shows the distribution of time instances  $T_\Theta$  and  $T_\Xi$  when the respective quantities  $\Theta_n$  and  
 362  $\Xi_n$  are maximum. These time instances agree well with the divergence times  $T_b$ . This confirms the single  
 363 trial finding in Fig. 7(A,B) that  $\Theta_n$  and  $\Xi_n$  are good markers for filter innovation divergence. Moreover  
 364 only few background members align and anti-align for  $r_l \leq r_H$  (small values of  $p_{a,u}$ ), whereas many more  
 365 background members align and anti-align for  $r_l \gg r_H$  (Fig. 8(B)). Conversely, each analysis member  
 366 aligns with its corresponding background member for  $r_l \leq r_H$  ( $q_a = 1$ ,  $q_u = 0$ ) and most analysis  
 367 members still align with their background members for  $r_l \gg r_H$  (Fig. 8(C)). This means that nonlocal  
 368 observations do poorly affect the direction of ensemble members in these cases.

### 369 Analytical description

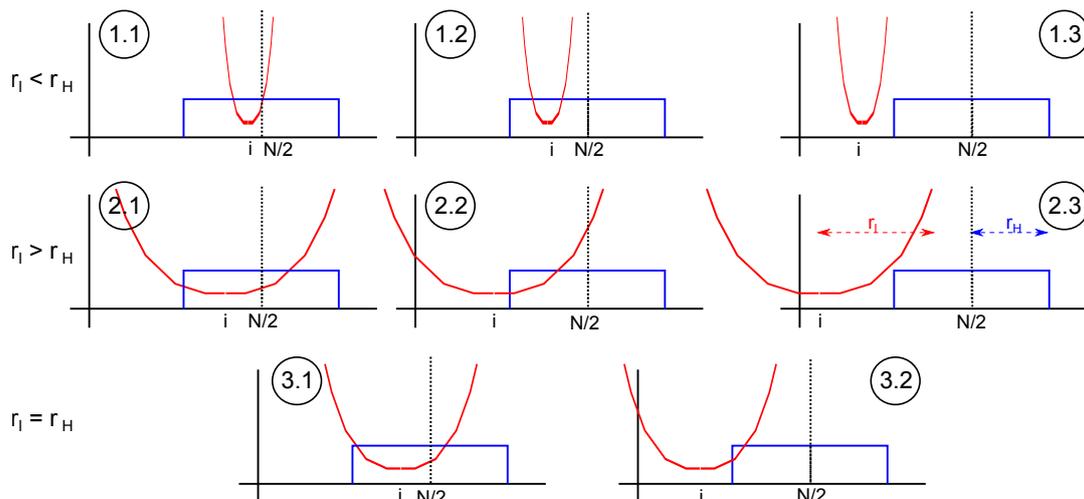
370 According to Fig. 9, there are different possible configurations of the sensitivity function with respect  
 371 to the localisation area. The localisation radius  $r_l$  may be smaller (cases 1) or larger (cases 2) than the  
 372 sensitivity width  $r_H$  or both may be equal (cases 3). In addition, it is insightful to distinguish observed and  
 373 unobserved grid points as already proposed in [18].

374 Now let us take a closer look at each case, cf. Fig. 9:



**Figure 8.** Divergence times and ensemble member alignment dependent on the localisation radius  $r_l$ . (A) Histogram of time of maximum  $\Theta_n$  ( $T_\Theta$ , black), time of maximum  $\Xi_n$  ( $T_\Xi$ , blue) and the divergence time  $T_b$  (green), see the Methods section 2.8 for definitions. (B) Histograms of alignment ratio  $p_a$  (black) and anti-alignment ratio  $p_u$  (red) defined in Eq. (24). (C) Histograms of alignment ratio  $q_a$  (black) and anti-alignment ratio  $q_u$  (red) defined in Eq. (26). In addition  $r_H = 5$  and results are based on the 200 numerical trials from Fig. 4.

375 • case **1.1**,  $r_l \leq r_H$ ,  $|i - \frac{N}{2}| \leq r_H$  and  
 376  $|i - \frac{N}{2}| \leq r_l$ : the localisation radius is smaller than the sensitivity function width and the observation  
 377 at spatial location  $N/2$  is located within the localisation radius about grid point  $i$ . Then, the analysis



**Figure 9.** Sketch of different configurations of sensitivity function and localisation area. The circles denote the different cases (n.m) The sensitivity function (blue) has its center at the center of the spatial domain and the localisation function (red) is located about model grid element  $i$ .

378 ensemble (20) and its mean (16) read

$$\begin{aligned}
 x_{o,i}^{a,l} &= \bar{x}_{o,i}^b + \mathbf{X}_{o,i} \mathbf{Y}^t \alpha_i + \mathbf{X}_{o,i} \mathbf{Y}^t \frac{\sqrt{\lambda_i}}{y^2} Y_l \\
 &+ \sum_{n=1}^{L-1} \mathbf{X}_{o,i} \mathbf{v}_n (\mathbf{v}_n)_l
 \end{aligned}
 \tag{27}$$

379 and

$$\bar{x}_{o,i}^a = \bar{x}_{o,i}^b + \mathbf{X}_{o,i} \mathbf{Y}^t \alpha_i
 \tag{28}$$

380 with the corresponding ensemble means at observed grid points  $\bar{x}_{o,i}^b$  and  $\bar{x}_{o,i}^a$ , the first guess  
 381 perturbations  $\mathbf{X}_{o,i}$  and the analysis ensemble members  $x_{o,i}^{a,l}$ .

382 • case **1.2**,  $r_l \leq r_H$ ,  $|i - \frac{N}{2}| \leq r_H$  and

383  $|i - \frac{N}{2}| > r_l$ : compared to case **1.1**, the grid point  $i$  is observed as well but the observation is outside  
 384 the localisation area; hence the analysis is identical to the first guess

$$\begin{aligned}
 \bar{x}_{o,i}^a &= \bar{x}_{o,i}^b \\
 x_{o,i}^{a,l} &= \bar{x}_{o,i}^b + (\mathbf{X})_{o,i}.
 \end{aligned}$$

385 • case **1.3**,  $r_l \leq r_H$ ,  $|i - N/2| > r_H$  and

386  $|i - N/2| > r_l$ : the grid point  $i$  is not observed and the observation is outside the localisation area  
 387 leading to

$$\begin{aligned}
 \bar{x}_{u,i}^a &= \bar{x}_{u,i}^b \\
 x_{u,i}^{a,l} &= \bar{x}_{u,i}^b + (\mathbf{X})_{u,i}
 \end{aligned}
 \tag{29}$$

388 with the corresponding unobserved ensemble means  $\bar{x}_{u,i}^b$  and  $\bar{x}_{u,i}^a$ , the unobserved ensemble  
 389 perturbations  $\mathbf{X}_{u,i}$  and the analysis ensemble member  $x_{u,i}^{a,l}$ .

390 • case **2.1**,  $r_l > r_H$ ,  $|i - \frac{N}{2}| \leq r_H$  and

391  $|i - \frac{N}{2}| \leq r_l$ : the localisation radius is larger than the sensitivity function width, the observation is  
 392 located within the localisation radius about the grid point  $i$  and all grid points are observed. This case  
 393 is equivalent to case **1.1** and the expressions for the analysis ensemble and mean hold as well.

394 • case **2.2**,  $r_l > r_H$ ,  $|i - \frac{N}{2}| > r_H$  and

395  $|i - \frac{N}{2}| \leq r_l$ : compared to case **2.1**, the observation is located within the localisation radius but grid  
 396 points are unobserved. Then

$$\bar{x}_{u,i}^a = \bar{x}_{u,i}^b + \mathbf{X}_{u,i} \mathbf{Y}^t \alpha_i
 \tag{30}$$

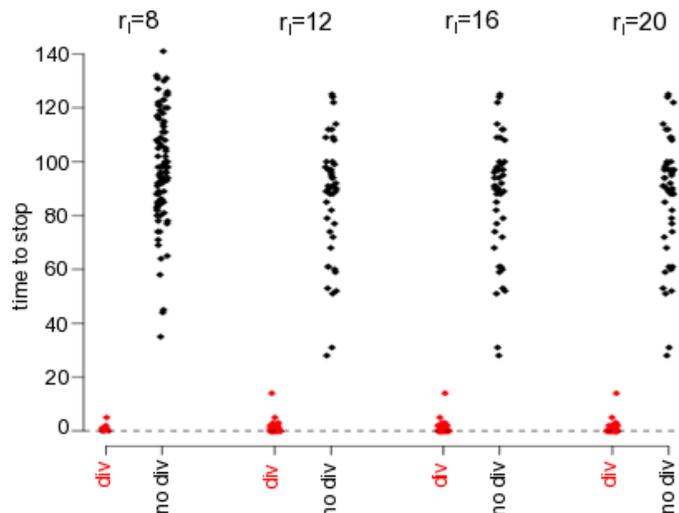
397 and

$$\begin{aligned}
 x_{u,i}^{a,l} &= \bar{x}_{u,i}^b + \mathbf{X}_{u,i} \mathbf{Y}^t \alpha_i + \mathbf{X}_{u,i} \mathbf{Y}^t \frac{\sqrt{\lambda_i}}{y^2} Y_l \\
 &\quad + \sum_{n=1}^{L-1} \mathbf{X}_{u,i} \mathbf{v}_n (\mathbf{v}_n)_l.
 \end{aligned} \tag{31}$$

- 398 • case **2.3**,  $r_l > r_H$ ,  $|i - \frac{N}{2}| > r_H$  and  
 399  $|i - \frac{N}{2}| > r_l$ : in this case, the grid points are unobserved and the observation is outside the localisation  
 400 area. Then the analysis is identical to the first guess and the case is equivalent to case **1.3**.
- 401 • case **3.1**,  $r_l = r_H$ ,  $|i - \frac{N}{2}| \leq r_H$ : the observation is located within the localisation radius about the  
 402 grid point  $i$ , the grid point is observed and the expressions in case **1.1** hold.
- 403 • case **3.2**,  $r_l = r_H$ ,  $|i - \frac{N}{2}| > r_H$ : the observation is not located within the localisation radius of grid  
 404 point  $i$ , then grid point is not observed and the expressions in case **1.3** hold.

405 Firstly, let us consider the limiting case of local observations with  $r_H = 1$ . Then case **1** does not exist.  
 406 This means that case **1** emerges for nonlocal observations only and Fig. 4 demonstrates that the filter  
 407 does not diverge for  $1 \leq r_l \leq 20$ . Moreover, the sensitivity function of the observation is non-zero at the  
 408 observation location only and hence the localisation of the observation to the position of the sensitivity  
 409 maximum (cf. Methods section 2.4) is trivial. In case **2**, this implies that updates at grid points far from  
 410 the observation location  $i \neq N/2$  consider the local observation with weighted observation error  $R_i$ . This  
 411 situation changes in case of nonlocal observations with  $r_H > 1$ . Then case **1** exists and analysis updates in  
 412 case **2** consider an erroneous estimate of the nonlocal observation at the single spatial location  $N/2$ . The  
 413 broader the sensitivity function and thus the larger  $r_H$ , the larger is the error induced by this localisation  
 414 approximation. Consequently, updates at grid points far from the observation location still consider the  
 415 observation with weighted observation error  $R_i$ , however the observation includes a much larger error than  
 416  $R_i$  introducing an analysis update error.

417 From a mathematical perspective, in cases **1.1**, **2.1** and **3.1** the LETKF updates observed grid points whereas  
 418 in the cases **1.3**, **2.3** and **3.2** no update is applied. These cases appear to be consistent since grid points that  
 419 contribute to the observation are updated by the observation and grid points that do not contribute to the  
 420 observation are not updated. Conversely, observed grid points in case **1.2** do not consider the observation  
 421 and are not updated although they contribute to the first guess in observation space. This missing update  
 422 contributes to the filter error and the filter divergence as stated in previous work [12]. Moreover, the  
 423 unobserved grid points in case **2.2** do consider the observation and are updated by the Kalman filter.  
 424 At a very first glance, this inconsistency may be detrimental similar to case **2.1**. However, it may be  
 425 arguable whether this inconsistency may contribute to the filter error. On the one hand, the background  
 426 error covariance propagates information between observed to unobserved grid points in each cycle step  
 427 and thus balances the missing contribution of the unobserved grid point to the observation. This may hold  
 428 true for system phenomena with a large characteristic spatial scale, such as wind advection or long-range  
 429 moisture transport in meteorology or, more generally, emerging long-range spatial synchronisation events.  
 430 However, on the other hand, if the background error covariance represents a bad estimate, e.g. due to  
 431 sampling errors or short-range synchronisation, the false (or inconsistent) update may enhance erroneous  
 432 propagated information and hence contributes to the filter divergence. This agrees with the vanishing  
 433 divergence in case of a full ensemble (cf. Fig. 5(bold dotted-dashed line)). Moreover, updates at unobserved  
 434 grid points may be erroneous due to model errors or the approximation error made by the reduction of



**Figure 10.** The divergence correlates with the weighted model-observation covariances at observed grid points  $A_o$ . The plots show the times of maxima  $T_o$  (cf. Eq. (35)) to stop, i.e.  $T_{\text{stop}} - T_o$ .  $T_o$  is the time when the mean model-observation covariance  $A_o$  is maximum, for divergent (red-colored with break time  $T_{\text{stop}}$ ) and non-divergent (black colored with maximum time  $T_{\text{stop}} = 200$ ) trials. Here it is  $r_H = 5$ .

435 nonlocal observations to a single location. The larger the localisation radius, the more distant are grid  
 436 points to the observation location and the less representative is the localised observation to distant grid  
 437 points.

438 Hence these two latter cases may cause detrimental effects. Consequently, cases 1 and 2, i.e.  $r_l \neq r_H$ ,  
 439 yields bad estimates of analysis updates that make the Kalman filter diverge. Conversely, case 3, i.e.  
 440  $r_l = r_H$ , involves consistent updates only and detrimental effects as described for the other cases are not  
 441 present. These effects may explain enhanced filter divergence for  $r_l \neq r_H$  and minimum filter divergence  
 442 for  $r_l = r_H$  seen in Fig. 3, and the minimum divergence rate at  $r_l \approx r_H$  shown in Fig. 4.

443 The important terms in case 2.2, i.e. Eqs. (30) and (31), are  $\mathbf{X}_{u,i}\mathbf{Y}^t$ ,  $\alpha_i$ ,  $\sqrt{\lambda_i}Y_l/y^2$ , and  
 444  $\sum_n \mathbf{X}_{u,i}\mathbf{v}_n(\mathbf{v}_n)_l$ . Equivalently, the missing terms in case 1.2 are  $\mathbf{X}_{o,i}\mathbf{Y}^t$ ,  $\alpha_i$ ,  $\sqrt{\lambda_i}Y_l/y^2$  and  
 445  $\sum_n \mathbf{X}_{o,i}\mathbf{v}_n(\mathbf{v}_n)_l$ . For instance,

$$\mathbf{c}_{o,u} = \mathbf{X}_{o,u}\mathbf{Y}^t = \sum_{l=1}^L (\mathbf{x}_{o,u}^{b,l} - \bar{\mathbf{x}}_{o,u}^b)(y^{b,l} - \bar{y}^b) \tag{32}$$

446 and  $\alpha_i$  appear in both cases 2.1 and 2.2. The terms  $\mathbf{c}_{o,u}$  represent the covariances between model and  
 447 observation perturbations over ensemble members and they may contribute differently to the intermittent  
 448 divergence with increasing  $|r_l - r_H|$ . For a closer investigation of these terms, let us consider

$$(\mathbf{c}_o)_i\alpha_i = \mathbf{X}_{o,i}\mathbf{Y}^t\alpha_i \tag{33}$$

449 in case 2.1 and

$$(\mathbf{c}_u)_i\alpha_i = \mathbf{X}_{u,i}\mathbf{Y}^t\alpha_i. \tag{34}$$

450 in case **2.2**. These terms represent the weighted ensemble covariances between model and observation  
 451 perturbations. To quantify their difference,

$$\begin{aligned} A_o &= \max_n \frac{1}{M_o} \sum_{i \in \mathcal{M}_o} (\mathbf{c}_o)_i(t_n) \alpha_i(t_n) \\ A_u &= \max_n \frac{1}{M_u} \sum_{i \in \mathcal{M}_u} (\mathbf{c}_u)_i(t_n) \alpha_i(t_n) \\ A &= A_o - A_u \end{aligned}$$

452 may be helpful. The term  $A_o$  ( $A_u$ ) is the maximum over time of the mean of  $(\mathbf{c}_o)_i \alpha_i$  ( $(\mathbf{c}_u)_i \alpha_i$ ). This mean is  
 453 computed over the set of observed (unobserved) grid points  $\mathcal{M}_o$  ( $\mathcal{M}_u$ ) with size  $M_o$  ( $M_u$ ). Consequently,  
 454  $A$  quantifies the difference of observed and unobserved **weighted** model-observation ensemble covariances,  
 455 while the unobserved covariances are down-weighted by  $\alpha_i$  compared to the observed covariances. This  
 456 down-weighting results from the fact that unobserved grid points are more distant from the observation  
 457 which yields smaller values of  $\alpha_i$ . **By definitions** (33) and (34), thus  $A < 0$  reflects **larger weighted**  
 458 **model-observation covariances** in unobserved than observed grid points.

459 The corresponding quantities

$$T_o = \arg \max_n \frac{1}{M_o} \sum_{i \in \mathcal{M}_o} (\mathbf{c}_o)_i(t_n) \alpha_i(t_n) \quad (35)$$

$$T_u = \arg \max_n \frac{1}{M_u} \sum_{i \in \mathcal{M}_u} (\mathbf{c}_u)_i(t_n) \alpha_i(t_n)$$

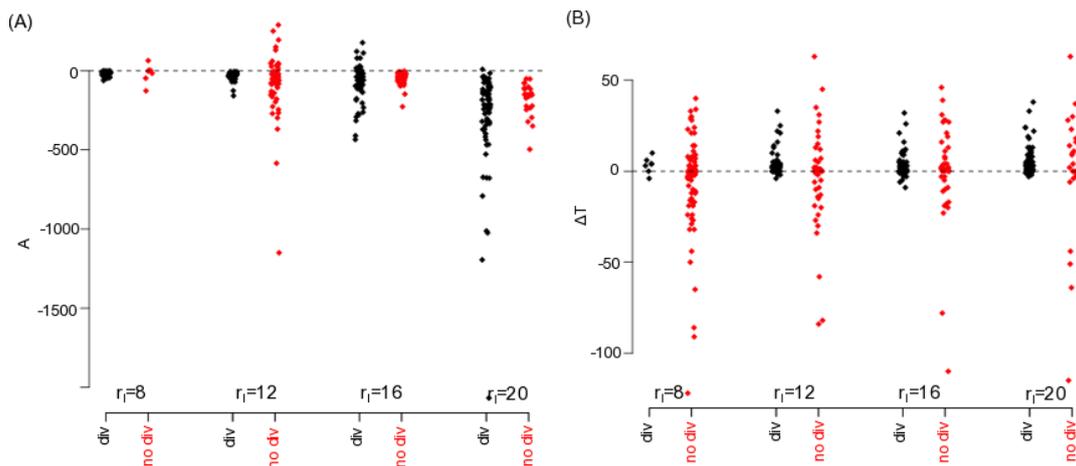
$$\Delta T = T_o - T_u \quad (36)$$

460 define the time instances when these maxima are reached and  $\Delta T$  is their difference. For instance, if  
 461  $\Delta T > 0$ , then the weighted model-observation covariances at observed grid points reach their maximum  
 462 before weighted model-observation covariances at unobserved grid points.

463 To illustrate the importance of  $A_o$  and its corresponding occurrence time  $T_o$ , Fig. 10 shows  $T_o$  relative to the  
 464 stop time  $T_{\text{stop}}$  of filter iteration, i.e.  $T_{\text{stop}} - T_o$ . For divergent trials,  $T_{\text{stop}} = T_b$  is the time of divergence  
 465 and for non-divergent trials  $T_{\text{stop}} = 200$  is the maximum time. Figure 10 reveals that  $T_o$  is very close to  
 466 the divergence time, whereas  $T_o$  is widely distributed about  $T_o = 110$  ( $T_{\text{stop}} - T_o = 90$ ) in non-divergent  
 467 trials. This indicates that  $A_o$  is strongly correlated with the underlying divergence mechanism.

468 Now that that  $A_o$  is strongly correlated with the filter innovation divergence, the question arises whether  
 469 the difference between weighted observed and unobserved model-observation covariances is related to the  
 470 innovation divergence. Figure 11 shows the distribution of  $A = A_o - A_u$  and  $\Delta T = T_o - T_u$  for **divergent**  
 471 **and non-divergent** experimental trials. Most trials exhibit stronger model-observation covariances in  
 472 unobserved grid points than in observed grid points ( $A < 0$ ), cf. Fig. 11(A), and the distribution variances  
 473 of divergent and **non-divergent** trials are significantly different (Fligner-Killeen test,  $p < 0.001$ ). Moreover,  
 474 the distribution of  $\Delta T$  in divergent trials is asymmetric since  $\Delta T > 0$  for almost all **divergent** trials  
 475 (see Fig. 11(B)). Hence **weighted** model-observation covariances in unobserved grid points reach their  
 476 maximum significantly earlier than in **observed** grid points. **Conversely the distribution of non-divergent**  
 477 **trials is more or less symmetric about  $\Delta T = 0$**  (Fligner-Killeen test,  $p < 0.0001$ ).

478 In this context, re-call that  $A_u > A_o$  but  $T_u < T_o$  in divergent trials, i.e. unobserved grid points reach their



**Figure 11.** Comparison of weighted model-observation covariances in observed and non-observed grid points. (A)  $A = A_o - A_u$  is the difference between **maximum weighted model-observation covariances** in observed and unobserved grid points. (B)  $\Delta T = T_o - T_u$  is the difference of times when the **weighted model-observation covariances** reach their maximum, cf. Eq. (36). It is  $r_H = 5$ .

479 larger maximum faster than observed grid points. This indicates that the model-observation covariance  $c_u$   
 480 reflects the instability of the filter.

### 481 3.2 Local and nonlocal observations

482 Several international weather services apply ensemble Kalman filters and assimilate both nonlocal  
 483 and local observations. Performing assimilation experiments similar to the experiments for nonlocal  
 484 observations but now with a single additional local observation at grid point  $i = N/2$  (cf. Methods section  
 485 2.7), the filter divergence rate  $\gamma$  indicates the filter stability. Figure 12 illustrates how local observations  
 486 affect the filter stability in addition to nonlocal observations. For  $r_H = 1$ , the filter diverges rarely due  
 487 to large innovations (with fewest trials at  $r_l \approx 10$ ) but at a larger number than in the absence of local  
 488 observations, cf. Fig. 4. Moreover, increasing the localisation radius yields a higher number of trials with  
 489 catastrophic filter divergence with a maximum catastrophic divergence rate at  $r_l \approx 10$ . In sum, the least  
 490 number of **divergent** trials occur at  $r_l = r_H = 1$  (blue curve in Fig. 12). A similar stability behavior occurs  
 491 for  $r_H = 5$  with a minimum innovation divergence rate at  $r_l \approx r_H$  and a maximum catastrophic divergence  
 492 rate at  $r_l \approx 10$ . Again, the least number of trials diverge at  $r_l = r_H$ .

493 Figure 1 motivates the present work demonstrating that nonlocal observations yield larger first guess  
 494 departures than for local observations only. Here, it is interesting to note that the numerical trial in Fig. 1  
 495 with nonlocal observations exceeds the innovation divergence threshold, cf. section 2.8, but has run  
 496 over all filter cycles for illustration reasons. Moreover, several trials with the same parameters exhibit  
 497 catastrophic filter divergence and the shown trial is a rare case. This divergence could have been avoided by  
 498 implementing stabilising features, such as ensemble enlargement [19], adaptive localisation [28], adaptive  
 499 inflation [18] or first guess check [13, 45]. However, these methods would have introduced additional  
 500 assimilation effects and the gained results would not have been comparable to findings and insights in the  
 501 remaining work.

## 4 DISCUSSION

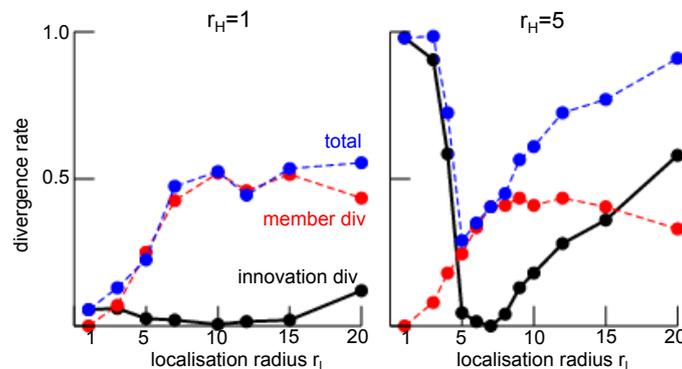
502 Ensemble Kalman filtering of nonlocal observations may increase the innovation in the filter process  
 503 leading to larger observation-background departure bias and RMSE, cf. Fig. 1. It is demanding to detect this  
 504 innovation divergence since it is finite and transient, i.e. of finite duration. At a first glance, this negative  
 505 impact is surprising since observations are thought to introduce additional knowledge to the system and  
 506 thus should improve forecasts or at least retain them. To understand better why nonlocal observations may  
 507 be detrimental, the present work performs numerical studies to identify markers of innovation divergence  
 508 and understand their origin.

### 509 Nonlocal observations facilitates filter divergence

510 The majority of previous stability studies of Kalman filtering involving nonlocal observations consider  
 511 catastrophic filter divergence. Kelly et al. [20] show analytically for a specific simple but non-trivial model  
 512 how catastrophic filter divergence of an ensemble Kalman filter is affected by nonlocal observations. The  
 513 work of Marx and Potthast [44] is an analytical discussion of the linear Kalman filter and the authors  
 514 derive corresponding stability conditions. Conversely, the present work considers intermittent innovation  
 515 divergence and, to our best knowledge, is one of the first to demonstrate this important effect numerically.  
 516 Intermittent innovation divergence is detrimental to forecasts and are visible, e.g., in first guess departure  
 517 statistics (Fig. 1). It occurs for a nonlocal observation only (Fig. 4) or for nonlocal and additional local  
 518 observation (Fig. 12). This holds true for almost all localisation radii.

### 519 Optimal localisation radius

520 Figures 4, 5, 6 and 12 show that innovation divergence depends on the relation between sensitivity  
 521 function width  $r_H$  and localisation radius  $r_l$ . The LETKF diverges least when  $r_l \approx r_H$  and hence this  
 522 choice of localisation radius is called optimal, i.e. the filter is least divergent. This insight agrees with the  
 523 finding in an operational weather prediction framework involving the LETKF [13]. The authors consider  
 524 an adaptive localisation for (nonlocal) satellite observations and choose the corresponding radius to the  
 525 sensitivity function width. In two different weather situations, this tight relation improves short- and  
 526 middle-range weather forecasts compared to the case of independent sensitivity width and localisation  
 527 radius. Figure 9 illustrates the possible reason for the detrimental effect of different sensitivity function  
 528 width and localisation radius: the LETKF is inconsistent if it updates the state at unobserved spatial



**Figure 12.** Rate of filter divergence  $\gamma$  (innovation divergence, black line) and catastrophic filter divergence (member divergence, red line) in the presence of a single local and a single nonlocal observation. The total number of **divergent** trials is the sum of innovation and member divergence-trials (blue line). Results are based on 200 numerical trials.

529 locations or does not update the state at observed spatial locations. Only if the sensitivity function and  
530 the localisation width are similar, then this detrimental effect is small. Such an inconsistency is in line  
531 with other inconsistencies in ensemble Kalman filters caused by localisation, cf. [46]. For instance, a full  
532 ensemble reduces inconsistencies for localisation radii larger than the sensitivity function width and yields  
533 filter stability (Fig. 5).

534 It is important to point out that, under certain conditions, it may be beneficial to further enlarge the  
535 localisation area compared to the sensitivity function. If the system's activity synchronizes on a larger  
536 spatial scale, then information is shared between observed and unobserved grid points and a larger  
537 localisation radius would be beneficial. Examples for such synchronisation phenomena are deep clouds  
538 or large-scale winds in meteorology or locally self-organised spots in physical complex systems. In other  
539 words, to decide how to choose the localisation radius one should take a closer look at the system's  
540 dynamics: if larger spatially synchronised phenomena are expected, then  $r_l \gg r_H$  is preferable, otherwise  
541  $r_l \approx r_H$ .

542 Several previous studies have derived optimal localisation radii for local observations in ensemble Kalman  
543 filter [47, 48, 49] and the specific LETKF [27, 50]. A variant of the LETKF localizes not in observation  
544 space as in the present work but in the spatial domain [34, 51, 31, 24], where the localisation of nonlocal  
545 observations has been studied as well [52]. There is the general agreement for local and non-local  
546 observations that the optimal localisation radius depends on the ensemble size and the observation error  
547 but seems to be independent on the model [50].

## 548 Origin of divergence

549 It is important to understand why some numerical trials diverge and some do not. Direct and indirect  
550 markers indicate which dynamical features play an important role in divergence. The most obvious direct  
551 markers are the absolute values of the innovation and the ensemble member perturbation spread  $\Theta_n$  and both  
552 increase sharply during filter innovation divergence, cf. Fig. 4, 6, 7(B), 8 and 12. Similarly, the covariation  
553 of observed and unobserved background errors  $\Xi_n$  also increases during divergence. Interestingly,  $\Theta_n$  and  
554  $\Xi_n$  remain finite and take their maxima just before the instance of divergence, cf. Fig. 8. The covariation  
555  $\Xi_n$  increases if both observed and unobserved errors increases. Kelly et al. [20] and Tong et al. [18] argue  
556 that this indicates a shift of power from observed to unobserved errors and that this shift is responsible  
557 for catastrophic divergence. The present findings support this line of argumentation and extends it to  
558 intermittent innovation divergence. This can be seen in Figure 11(A). It shows larger mean weighted  
559 model-observation error covariances (i.e. ensemble error covariances) in unobserved grid points than  
560 in observed grid points ( $A < 0$ ) and these weighted model-observation covariances increase faster in  
561 unobserved grid points than in observed grid points. In addition, the larger the localisation radius  $r_l > r_H$ ,  
562 the larger the ensemble error in unobserved grid points compared to observed grid points. Hence the  
563 model-observation covariance reflects a degree of instability (and thus of divergence) in the filter andn this  
564 is stronger in unobserved grid points than in observed grid points.

565 Figures 4, 5, 6 and 12 provide further evidence on possible error sources that yield filter divergence. The  
566 asymmetry of the divergence rates with respect to  $r_l \approx r_H$  hints different underlying filter divergence  
567 contributions. If  $r_l < r_H$ , too few grid points are updated by the nonlocal observation (Fig. 9) although  
568 they are observed. Consequently observations include contributions from non-updated grid points which  
569 might yield large observation contributions from large model magnitudes and hence this error mechanism  
570 is rather strong. Fertig et al. [12] have identified this case as a possible source of divergence and proposed  
571 to adapt the localisation radius to the sensitivity function width. In fact, this removes case 1 in Fig. 9 and  
572 stabilises the filter for  $r_l < r_H$ .

573 For  $r_l \gg r_H$ , a large number of grid points are updated which, however, consider an observation with  
574 a large intrinsic error **resulting from, e.g., a too low number of ensemble members**. The corresponding  
575 assimilation error is more subtle than for  $r_l < r_H$  and increases for larger localisation radii only. The  
576 localised nonlocal observation comprises a representation error due to the reduction of the broad sensitivity  
577 function to a single location. **For small ensembles, this implicit observation error contributes to the analysis**  
578 **update error and, finally, to filter divergence.** In sum, the two inconsistencies illustrated in Fig. 9 and  
579 derived in section 3.1 represent two possible contributions to the filter divergence **for a low number of**  
580 **ensemble members. Conversely, for a full ensemble, intrinsic error contributions are well reduced rendering**  
581 **the filter more stable (Fig. 5).**

582 Moreover, there is some evidence that ensemble member alignment may cause catastrophic filter  
583 divergence [21, 19, 20]. Figure 8 shows such indirect markers indicating weak member anti-alignment for  
584  $r_l \leq r_H$  but enhanced alignment and anti-alignment for  $r_l > r_H$ . The authors in [19] argue that finite  
585 ensemble sizes cause the ensemble to align in case of divergence and Ng et al. [53] show that the ensemble  
586 members may align with the most unstable phase space direction. However, our results reveal that member  
587 alignment does not represent the major mechanism for innovation divergence. Conversely, Fig. 8 provides  
588 evidence for alignment of analysis increments and background perturbations when the filter diverges. This  
589 alignment indicates that the analysis members point into the same direction as the background members.  
590 For instance, if background member perturbations point to less stable locations in phase space, then the  
591 LETKF does not correct this direction and the new analysis state is less stable, cf. the model example  
592 in [20]. This shows accordance to the reasoning in Ng et al. [53].

593 In addition to the alignment mechanism, Eq. (32) represents the covariation of ensemble perturbations in  
594 spatial and observation space at observed and unobserved spatial locations. For observed spatial locations,  
595 it is maximum just before the innovation divergence time. Moreover, it reaches its maximum at unobserved  
596 locations almost always before the maximum at observed locations are reached (Fig. 11). It seems this new  
597 feature represent an important contribution to the innovation divergence and future work will analyse this  
598 covariation in more detail.

## 599 **Limits and outlook**

600 The present work considers the specific case of finite low ensemble size and application of the localisation  
601 scheme. To understand better the origin of the filter divergence, it is insightful to study in detail the limiting  
602 case of large ensemble sizes, i.e. close to the model dimension, and a neglect of localisation. Although this  
603 limit is far from practice in seismology and meteorology, where the model systems are too large to study  
604 this limit, nevertheless this limit study is of theoretical interest and future work will consider it in some  
605 detail.

606 There is some evidence that the optimal localisation radius is flow-dependent [54, 55], whereas we assume  
607 a constant radius. In addition, the constrained choice of parameters and missing standard techniques to  
608 prevent divergence, such as adaptive inflation and adaptive observation error, limits the present work  
609 in generality and interpretation and thus makes it hard to derive decisive conclusions. Future work will  
610 implement adaptive schemes [56, 57] in a more realistic model framework.

611 In the majority of studies, the present work considers a non-smooth boxcar sensitivity function in order  
612 to distinguish observed and unobserved grid points. Although this simplification allows to gain deeper  
613 understanding of possible contributions to the filter divergence, the sensitivity function is unrealistic. A  
614 more realistic sensitivity function is smooth and unimodal or bimodal. Figure 5 shows that such a sensitivity  
615 function yields a minimum divergence rate but the localisation radius at the minimum rate is much smaller  
616 than the sensitivity function width. Consequently, the line of argumentation about Fig. 9 does not apply

617 here since there is no clear distinction of observed and unobserved grid points anymore. Future work will  
618 attempt to consider smooth unimodal sensitivity functions.  
619 Moreover, the localisation scheme of nonlocal observations applied in the present work is very basic due to  
620 its choice of the maximum sensitivity as the observations location. Future work will investigate cut-off  
621 criteria as such in [12] that chooses the location of nonlocal observations in the range of the sensitivity  
622 function. Fertig et al. [12] also have shown that such a cut-off criterion improves first guess departure  
623 statistics and well reduces the divergence for localisation radii that are smaller than the sensitivity weighting  
624 function.  
625 Nevertheless the present work introduces the problem of intermittent innovation divergence, extends lines  
626 of reason on the origin of filter divergence to nonlocal observation and proposes new markers of innovation  
627 divergence.  
628

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## REFERENCES

- 632 [1] Bengtsson L, Ghil M, Källén E, editors. *Dynamic Meteorology: Data Assimilation Methods, Applied*  
633 *Mathematical Sciences*, vol. 36 (Springer) (1981).
- 634 [2] Luo X, Bhakta T, Jakobsen M, Navdal G. Efficient big data assimilation through sparse representation:  
635 A 3d benchmark case study in petroleum engineering. *PLoS One* **13** (2018) e0198586. doi:10.1371/  
636 journal.pone.0198586.
- 637 [3] Hutt A, Stannat W, Potthast R, editors. *Data Assimilation and Control: Theory and Applications in*  
638 *Life Sciences* (Frontiers Media) (2019). doi:10.3389/978-2-88945-985-8.
- 639 [4] Schiff SJ. *Neural Control Engineering* (Cambridge, MA: MIT Press) (2011).
- 640 [5] Nakamura G, Potthast R. *Inverse Modeling*. 2053-2563 (IOP Publishing) (2015). doi:10.1088/  
641 978-0-7503-1218-9.
- 642 [6] Asch M, Bocquet M, Nodet M. *Data Assimilation: Methods, Algorithms, and Applications*  
643 (Philadelphia: SIAM) (2016).
- 644 [7] Hunt B, Kostelich E, Szunyogh I. Efficient data assimilation for spatiotemporal chaos: A local  
645 ensemble transform Kalman filter. *Physica D* **230** (2007) 112–126.
- 646 [8] Schraff C, Reich H, Rhodin A, Schomburg A, Stephan K, Perianez A, et al. Kilometre-scale ensemble  
647 data assimilation for the cosmo model (kenda). *Q. J. R. Meteorol. Soc.* **142** (2016) 1453–1472.  
648 doi:10.1002/qj.2748.
- 649 [9] Schomburg A, Schraff C, Potthast R. A concept for the assimilation of satellite cloud information  
650 in an ensemble Kalman filter: single-observation experiments. *Q. J. R. Meteorol. Soc.* **141** (2015)  
651 893–908. doi:10.1002/qj.2748.
- 652 [10] Miyoshi T, Sato Y. Assimilating satellite radiances with a local ensemble transform Kalman filter  
653 (letkf) applied to the jma global model (gsm). *SOLA* **3** (2007) 37–40. doi:10.2151/sola.2007-010.
- 654 [11] Kurzrock F, Cros S, Ming F, Otkin J, Hutt A, Linguet L, et al. A review of the use of geostationary  
655 satellite observations in regional-scale models for short-term cloud forecasting. *Meteorologische*  
656 *Zeitschrift* **27** (2018) 277–298. doi:10.1127/metz/2018/0904.

- 657 [12] Fertig EJ, Hunt BR, Ott E, Szunyogh I. Assimilating non-local observations with a local ensemble  
658 Kalman filter. *Tellus A* **59** (2007) 719–730. doi:10.1111/j.1600-0870.2007.00260.x.
- 659 [13] Hutt A, Schraff C, Anlauf H, Bach L, Baldauf M, Bauernschubert E, et al. Assimilation of SEVIRI  
660 water vapour channels with an ensemble Kalman filter on the convective scale. *Front. Earth Sci.* **8**  
661 (2019) 70. doi:10.3389/feart.2020.00070.
- 662 [14] Furrer R, Bengtsson T. Estimation of high-dimensional prior and posterior covariance matrices in  
663 Kalman filter variants. *J. Multivar. Ana.* **98** (2007) 227–255.
- 664 [15] Anderson JL. An ensemble adjustment Kalman filter for data assimilation. *Mon. Wea. Rev.* **129**  
665 (2001) 2884–2903.
- 666 [16] Hamill TM, Whitaker JS, Snyder C. Distance-dependent filtering of background error covariance  
667 estimates in an ensemble Kalman filter. *Mon. Wea. Rev.* **129** (2001) 2776–2790. doi:10.1175/  
668 1520-0493.
- 669 [17] Tong XT, Majda A, Kelly D. Nonlinear stability and ergodicity of ensemble based Kalman filters.  
670 *Nonlinearity* **29** (2016) 657–691. doi:10.1088/0951-7715/29/2/657.
- 671 [18] Tong XT, Majda A, Kelly D. Nonlinear stability of ensemble Kalman filters with adaptive covariance  
672 inflation. *Commun. Math. Sci.* **14** (2016) 1283–1313. doi:10.4310/cms.2016.v14.n5.a5.
- 673 [19] Gottwald G, Majda AJ. A mechanism for catastrophic filter divergence in data assimilation for sparse  
674 observation networks. *Nonlin. Processes Geophys.* **20** (2013) 705–712. doi:10.5194/npg-20-705-2013.
- 675 [20] Kelly D, Majda A, Tong XT. Concrete ensemble Kalman filters with rigorous catastrophic filter  
676 divergence. *Proc. Natl. Acad. Sci. USA.* **112** (2015) 10589–10594. doi:10.1073/pnas.1511063112.
- 677 [21] Majda A, Harlim J. Catastrophic filter divergence in filtering nonlinear dissipative systems. *Comm.*  
678 *Math. Sci.* **8** (2008) 27–43.
- 679 [22] Migliorini S, Candy B. All-sky satellite data assimilation of microwave temperature sounding channels  
680 at the met office. *Quart. J. Roy. Meteor. Soc.* **145** (2019) 867–883. doi:10.1002/qj.3470.
- 681 [23] Lorenz EN, Emanuel KA. Optimal sites for supplementary weather observations: Simulations  
682 with a small model. *J. Atmos. Sci.* **555** (1998) 399–414. doi:10.1175/1520-0469(1998)055<0399:  
683 SFSWO>2.0.CO;2.
- 684 [24] Bishop CH, Whitaker JS, Lei L. Gain form of the Ensemble Transform Kalman Filter and its relevance  
685 to satellite data assimilation with model space ensemble covariance localization. *Mon. Wea. Rev.* **145**  
686 (2017) 4575–4592. doi:10.1175/MWR-D-17-0102.1.
- 687 [25] Waller JA, Dance SL, Lawless AS, Nichols NK. Estimating correlated observation error statistics  
688 using an ensemble transform Kalman filter. *Tellus A* **66** (2014) 23294. doi:10.3402/tellusa.v66.23294.
- 689 [26] Houtekamer PL, Zhang F. Review of the ensemble Kalman filter for atmospheric data assimilation.  
690 *Mon. Wea. Rev.* **144** (2016) 4489–4532. doi:10.1175/MWR-D-15-0440.1.
- 691 [27] Perianez A, Reich H, Potthast R. Optimal localization for ensemble Kalman filter systems. *J. Met.*  
692 *Soc. Japan* **92** (2014) 585–597. doi:10.2151/jmsj.2014-605.
- 693 [28] Greybush SJ, Kalnay E, Miyoshi T, Ide K, Hunt BR. Balance and ensemble Kalman filter localization  
694 techniques. *Mon. Wea. Rev.* **139** (2011) 511–522. doi:10.1175/2010MWR3328.1.
- 695 [29] Gaspari G, Cohn S. Construction of correlation functions in two and three dimensions. *Q. J. R. Meteo.*  
696 *Soc.* **125** (1999) 723–757.
- 697 [30] Nadeem A, Potthast R. Transformed and generalized localization for ensemble methods in data  
698 assimilation. *Math. Meth. Appl. Sci.* **39** (2016) 619–634. doi:10.1002/mma.3496.
- 699 [31] Bishop CH, Hodyss D. Ensemble covariances adaptively localized with eco-rap. part 2: a strategy for  
700 the atmosphere. *Tellus* **61A** (2009) 97–111. doi:10.1111/j.1600-0870.2008.00372.

- 701 [32] Leng H, Song J, Lu F, Cao X. A new data assimilation scheme: the space-expanded ensemble  
702 localization Kalman filter. *Adv. Meteorol.* **2013** (2013) 410812. doi:10.1155/2013/410812.
- 703 [33] Miyoshi T, Yamane S. Local Ensemble Transform Kalman Filtering with an AGCM at a T159/L48  
704 resolution. *Mon. Wea. Rev.* **135** (2007) 3841–3861. doi:10.1175/2007MWR1873.1.
- 705 [34] Farchi A, Boquet M. On the efficiency of covariance localisation of the ensemble Kalman filter using  
706 augmented ensembles. *Front. Appl. Math. Stat.* **5** (2019) 3. doi:10.3389/fams.2019.00003.
- 707 [35] Lei L, Whitaker JS. Model space localization is not always better than observation space  
708 localization for assimilation of satellite radiances. *Mon. Wea. Rev.* **143** (2015) 3948–3955.  
709 doi:10.1175/MWR-D-14-00413.1.
- 710 [36] Campbell WF, Bishop CH, Hodyss D. Vertical covariance localization for satellite radiances in  
711 ensemble Kalman filters. *Mon. Wea. Rev.* **138** (2010) 282–290. doi:10.1175/MWR3017.1.
- 712 [37] Houtekamer PL, Mitchell H, Pellerin G, Buehner M, Charron M, Spacek L, et al. Atmospheric data  
713 assimilation with an ensemble Kalman filter: Results with real observations. *Mon. Wea. Rev.* **133**  
714 (2005) 604–620. doi:10.1175/MWR-2864.1.
- 715 [38] Higham NJ. *Accuracy and stability of numerical algorithms* (SIAM), 2nd edn. (2002).
- 716 [39] Anderson JL, Anderson SL. A monte carlo implementation of the nonlinear filtering problem to  
717 produce ensemble assimilations and forecasts. *Mon. Wea. Rev.* **127** (1999) 2741–2758. doi:10.1175/  
718 1520-0493.
- 719 [40] Luo X, Hoteit I. Covariance inflation in the ensemble Kalman filter: A residual nudging perspective  
720 and some implications. *Mon. Weath. Rev.* **141** (2013) 3360–3368. doi:10.1175/MWR-D-13-00067.1.
- 721 [41] Hamill TM, Whitaker JS. What constrains spread growth in forecasts initialized from ensemble  
722 Kalman filters ? *Mon. Wea. Rev.* **139** (2011) 117–131. doi:10.1175/2010MWR3246.1.
- 723 [42] Mitchell HL, Houtekamer PL. An adaptive ensemble Kalman filter. *Mon. Weath. Rev.* **128** (2000)  
724 416–433. doi:10.1175/1520-0493.
- 725 [43] Grewal MS, Andrews AP. *Kalman filtering: Theory and practice using MATLAB* (John Wiley &  
726 Sons), 2nd edn. (2001).
- 727 [44] Marx B, Potthast R. On instabilities in data assimilation algorithms. *Mathematics* **8** (2012) 27–43.  
728 doi:10.1007/s13137-012-0034-5.
- 729 [45] Lahoz WA, Schneider P. Data assimilation: making sense of earth observation. *Front. Environ. Sci.* **2**  
730 (2014) 16. doi:10.3389/fenvs.2014.00016.
- 731 [46] Tong XT. Performance analysis of local ensemble Kalman filter. *J. Nonlinear. Sci.* **28** (2018)  
732 1397–1442. doi:10.1007/s00332-018-9453-2.
- 733 [47] Ying Y, Zhang F, Anderson J. On the selection of localization radius in ensemble filtering for multiscale  
734 quasigeostrophic dynamics. *Mon. Wea. Rev.* **146** (2018) 543–560. doi:10.1175/MWR-D-17-0336.1.
- 735 [48] Miyoshi T, Kondo K. A multi-scale localization approach to an ensemble Kalman filter. *SOLA* **9**  
736 (2013) 170–173.
- 737 [49] Migliorini S. Information-based data selection for ensemble data assimilation. *Quart. J. Roy. Meteor.*  
738 *Soc.* **139** (2013) 2033–2054.
- 739 [50] Kirchgessner P, Nerger L, Bunse-Gerstner A. On the choice of an optimal localization radius  
740 in Ensemble Kalman Filter methods. *Mon. Wea. Rev.* **142** (2014) 2165–2175. doi:10.1175/  
741 MWR-D-13-00246.1.
- 742 [51] Bishop CH, Whitaker JS, Lei L. Commentary: On the efficiency of covariance localisation of the  
743 ensemble Kalman filter using augmented ensembles by alban farchi and marc bocquet. *Front. Appl.*  
744 *Math. Stat.* **accepted** (2020).

- 745 [52] Lei L, Whitaker JS, Bishop C. Improving assimilation of radiance observations by implementing model  
746 space localization in an ensemble Kalman filter. *J. Adv. Model. Earth Syst.* **10** (2018) 3221–3232.  
747 doi:10.1029/2018MS001468.
- 748 [53] Ng GHC, McLaughlin D, Entekhabi D, Ahanin A. The role of model dynamics in ensemble Kalman  
749 filter performance for chaotic systems. *Tellus A* **63** (2011) 958–977.
- 750 [54] Zhen Y, Zhang F. A probabilistic approach to adaptive covariance localization for serial ensemble  
751 square root filters. *Mon. Wea. Rev.* **142** (2014) 4499–4518.
- 752 [55] Flowerdew J. Towards a theory of optimal localisation. *Tellus A* **67** (2015) 25257.
- 753 [56] Lee Y, Majda AJ, Qi D. Preventing catastrophic filter divergence using adaptive additive inflation for  
754 baroclinic turbulence. *Mon. Wea. Rev.* **145** (2017) 669–682. doi:10.1175/MWR-D-16-0121.1.
- 755 [57] Miyoshi T. The gaussian approach to adaptive covariance inflation and its implementation with  
756 the local ensemble transform Kalman filter. *Mon. Wea. Rev.* **139** (2011) 1519–1535. doi:10.1175/  
757 2010MWR3570.1.